

MicroBooNE:  
+ Neutrino Physics at the  
Dawn of the Liquid Argon TPC Era

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Weak Interactions Discussion Group Seminar @ Yale  
November 11, 2013



# No sterile neutrinos in the Standard Model\*

*\*Minimally* extended to account for neutrino mass

# Three-Neutrino Oscillation Parameters

**Ingredients for neutrino oscillation:**

**neutrinos have non-zero masses**

**+**

**(neutral) leptons mix**

weak (“flavor”) states

“mass” states

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3}e^{i\delta} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

3×3 unitary mixing matrix  $U$

# Three-Neutrino Oscillation Parameters

1. Why 3 “flavor” states?
2. Why 3 “mass” states?

weak (“flavor”) states

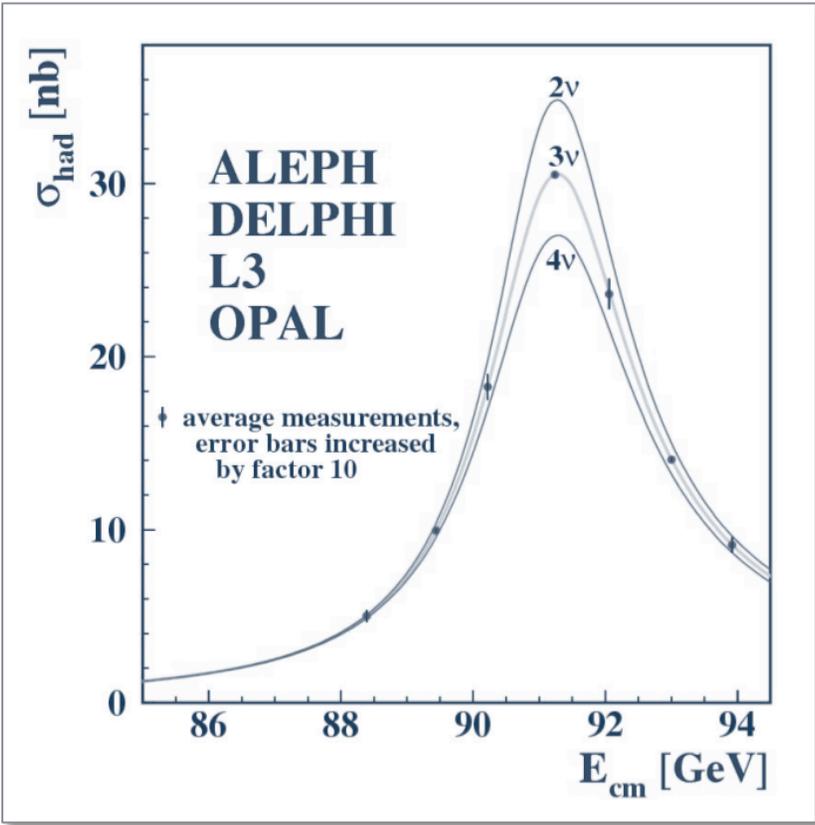
“mass” states

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3}e^{i\delta} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

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# Three-Neutrino Oscillation Parameters

## 1. Why 3 “flavor” states?



$$N_\nu = \frac{\Gamma_{\text{inv}}}{\Gamma_\ell} \left( \frac{\Gamma_\ell}{\Gamma_\nu} \right)_{\text{SM}}$$

$$= 2.984 \pm 0.008$$

[Phys. Reports 427, 257 (2006)]

Measurement of the invisible Z width:  $Z \rightarrow \nu\bar{\nu}$

# Three-Neutrino Oscillation Parameters

## 2. Why 3 “mass” states?

### 1. Theoretical prejudice

### 2. Limits on number of light neutrino species

from cosmology\*

Model	Data	$N_{eff}$	
$N_{eff}$	W-5+BAO+SN+ $H_0$	$4.13^{+0.87(+1.76)}_{-0.85(-1.63)}$	
	W-5+LRG+ $H_0$	$4.16^{+0.76(+1.60)}_{-0.77(-1.43)}$	
	W-5+CMB+BAO+XLF+ $f_{gas}+H_0$	$3.4^{+0.6}_{-0.5}$	
	W-5+LRG+maxBCG+ $H_0$	$3.77^{+0.67(+1.37)}_{-0.67(-1.24)}$	
	W-7+BAO+ $H_0$	$4.34^{+0.86}_{-0.88}$	
	W-7+LRG+ $H_0$	$4.25^{+0.76}_{-0.80}$	
	W-7+ACT	$5.3 \pm 1.3$	
	W-7+ACT+BAO+ $H_0$	$4.56 \pm 0.75$	
	W-7+SPT	$3.85 \pm 0.62$	
	W-7+SPT+BAO+ $H_0$	$3.85 \pm 0.42$	
	W-7+ACT+SPT+LRG+ $H_0$	$4.08^{(+0.71)}_{(-0.68)}$	
	W-7+ACT+SPT+BAO+ $H_0$	$3.89 \pm 0.41$	
	$N_{eff}+f_\nu$	W-7+CMB+BAO+ $H_0$	$4.47^{(+1.82)}_{(-1.74)}$
		W-7+CMB+LRG+ $H_0$	$4.87^{(+1.86)}_{(-1.75)}$
$N_{eff}+\Omega_k$	W-7+BAO+ $H_0$	$4.61 \pm 0.96$	
	W-7+ACT+SPT+BAO+ $H_0$	$4.03 \pm 0.45$	
$N_{eff}+\Omega_k+f_\nu$	W-7+ACT+SPT+BAO+ $H_0$	$4.00 \pm 0.43$	
$N_{eff}+f_\nu+w$	W-7+CMB+BAO+ $H_0$	$3.68^{(+1.90)}_{(-1.84)}$	
	W-7+CMB+LRG+ $H_0$	$4.87^{(+2.02)}_{(-2.02)}$	
$N_{eff}+\Omega_k+f_\nu+w$	W-7+CMB+BAO+SN+ $H_0$	$4.2^{+1.10(+2.00)}_{-0.61(-1.14)}$	
	W-7+CMB+LRG+SN+ $H_0$	$4.3^{+1.40(+2.30)}_{-0.54(-1.09)}$	

[pre-Planck data: arXiv:1204.5379]

# Three-Neutrino Oscillation Parameters

Mixing matrix parameterization for two-mass-scale dominance scenario:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3}e^{i\delta} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



three mixing angles:

$$U = \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}$$

“solar”  
 $\theta_{12} \approx 34^\circ$

“reactor”  
 $\theta_{13} = 9^\circ$

“atmospheric”  
 $\theta_{23} \approx 45^\circ$

and a CP-violating phase:

If  $\delta \neq 0$ , then have CP violation  $\Rightarrow P(\nu_\mu \rightarrow \nu_e) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

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three mixing angles:

*As of 2012, the only unknown mixing parameter!*

$$U = \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}$$

“solar”  
 $\theta_{12} \approx 34^\circ$

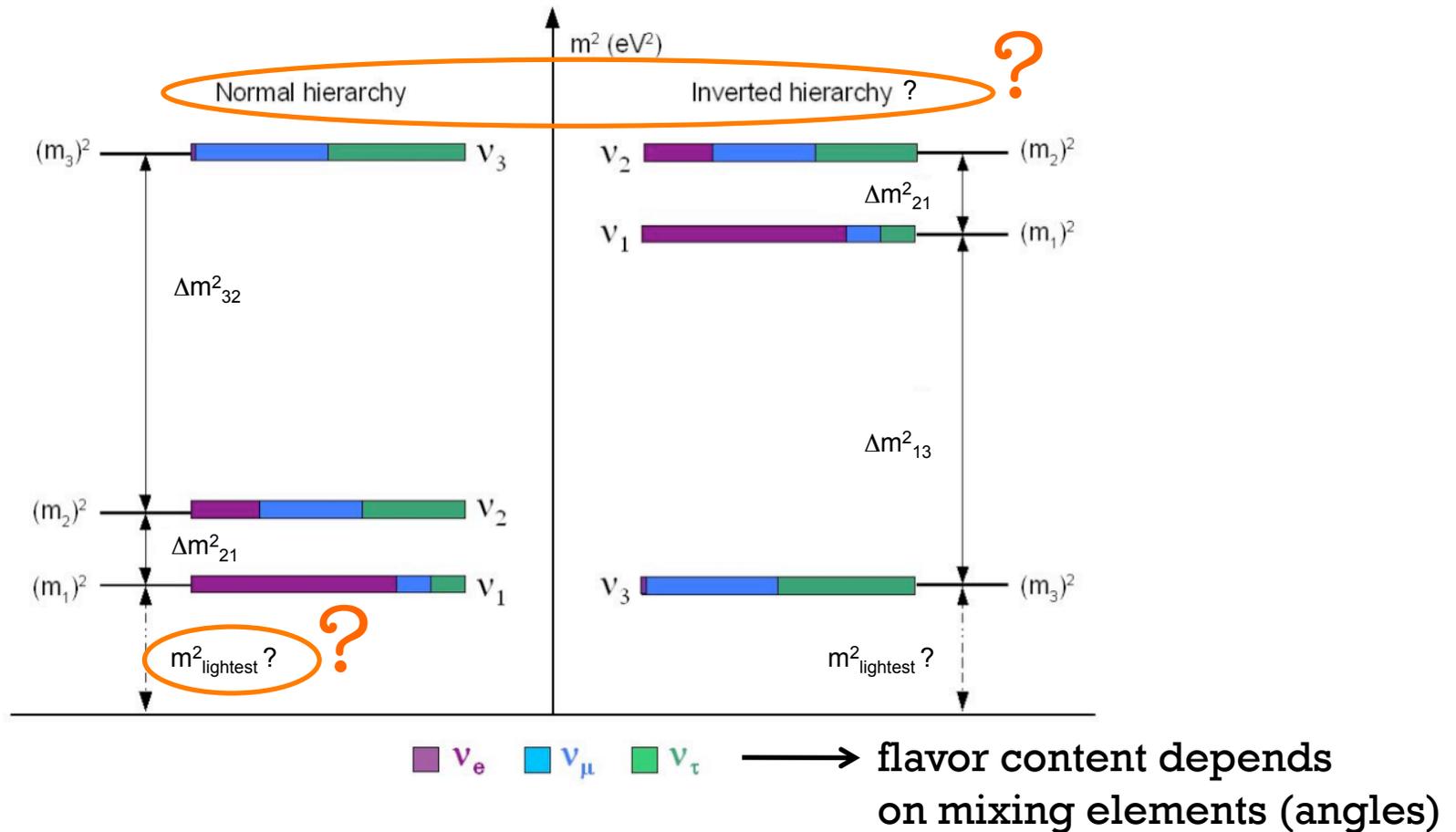
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# Three-Neutrino Oscillation Parameters

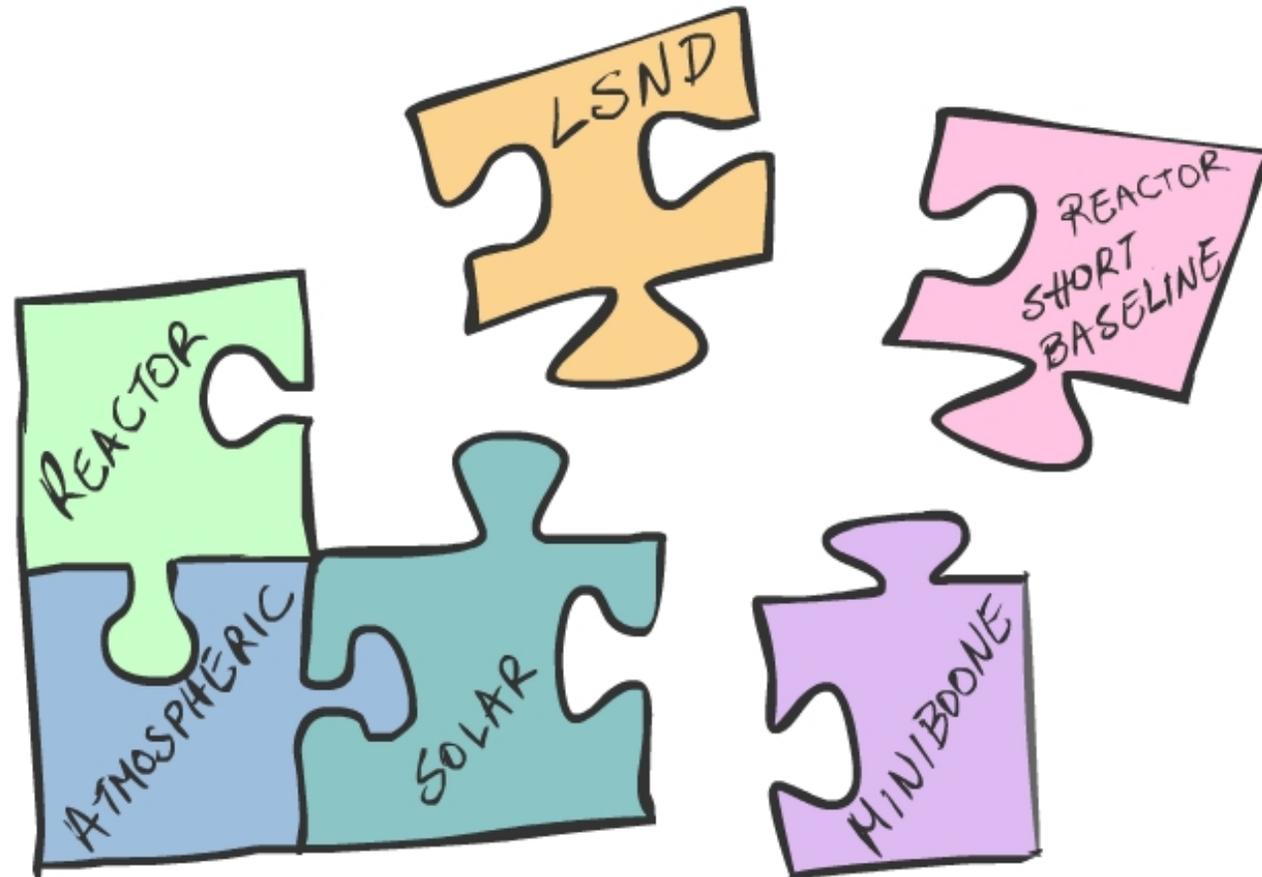


Three mass splittings:

$$\Delta m_{21}^2 = m_2^2 - m_1^2 \quad \Delta m_{32}^2 = m_3^2 - m_2^2 \quad \Delta m_{31}^2 = m_3^2 - m_1^2$$

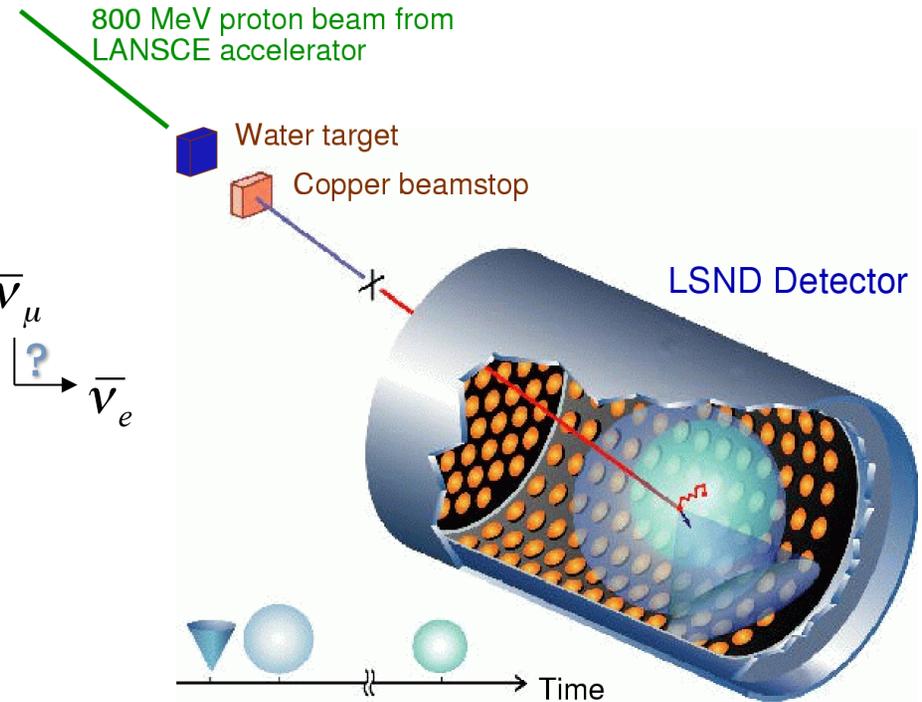
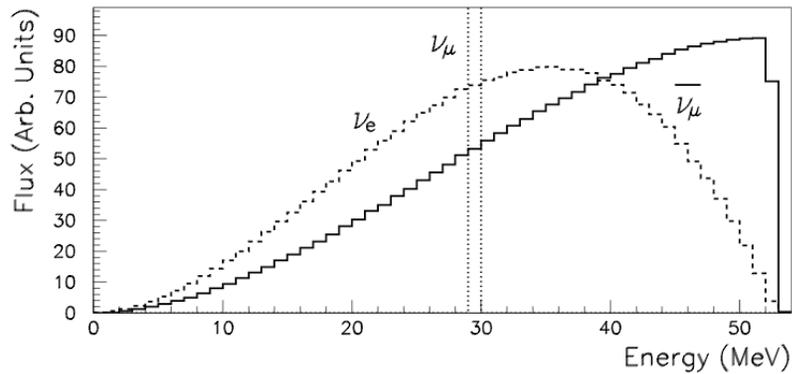
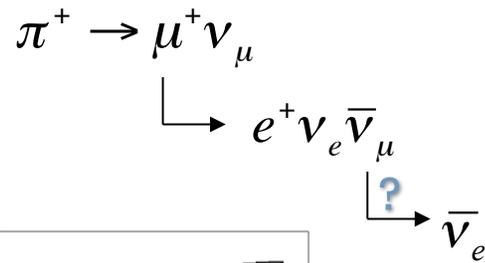
but only two are independent since the framework assumes only three mass states.

# Why or why not sterile neutrinos?



# Puzzle piece #1: LSND Experiment

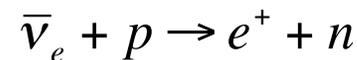
$\mu^+$  decay-at-rest experiment:



Well-predicted neutrino flux and cross-section.

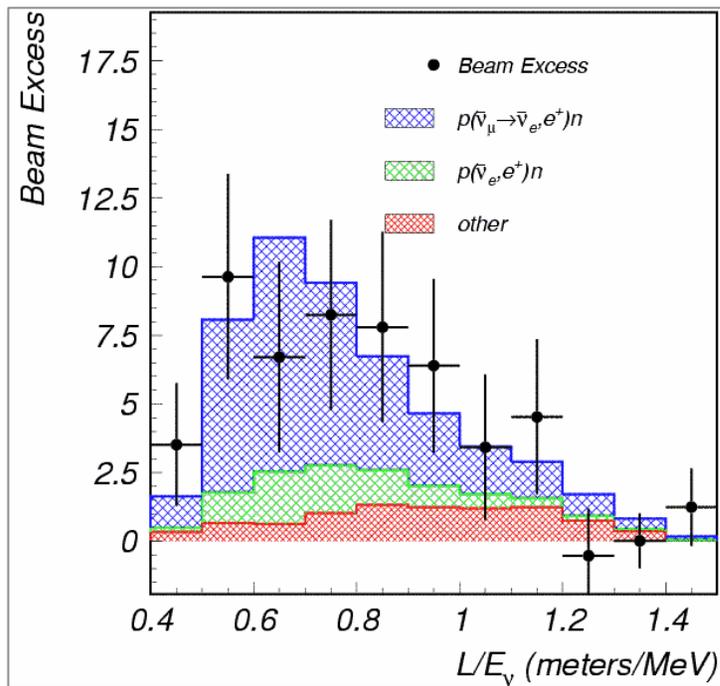
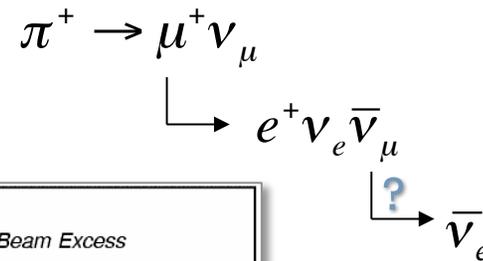
Very low  $\bar{\nu}_e$  backgrounds.

$\bar{\nu}_e$  detection via inverse-beta-decay:  
(coincidence signal)



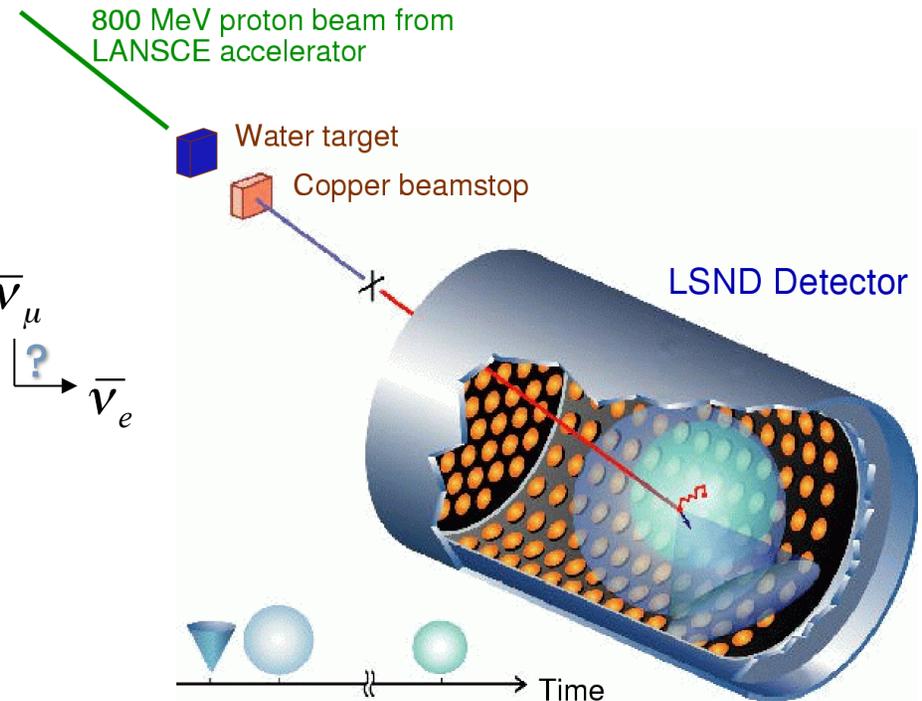
# Puzzle piece #1: LSND Experiment

$\mu^+$  decay-at-rest experiment:



Observed excess of  $\bar{\nu}_e$   
described by oscillation probability:  
 $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = (0.264 \pm 0.067 \pm 0.045) \%$

(3.8 $\sigma$  evidence)

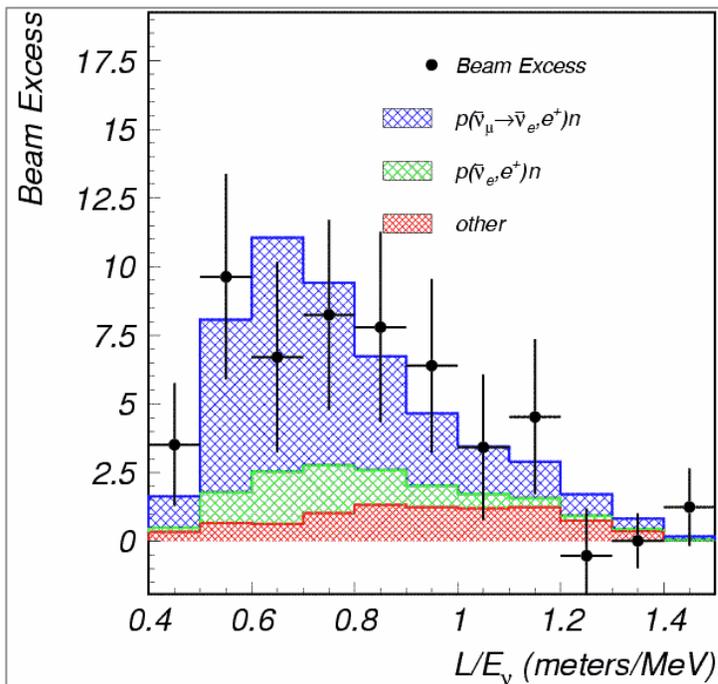


[C. Athanassopoulos et al., Phys. Rev. Lett. 75, 2650 (1995);  
81,1774(1998); A.Aguilaretal., Phys. Rev. D64, 112007(2001).]

# Puzzle piece #1: LSND Experiment

Points to large  $\Delta m^2$   
if interpreted as  
two-neutrino oscillations:

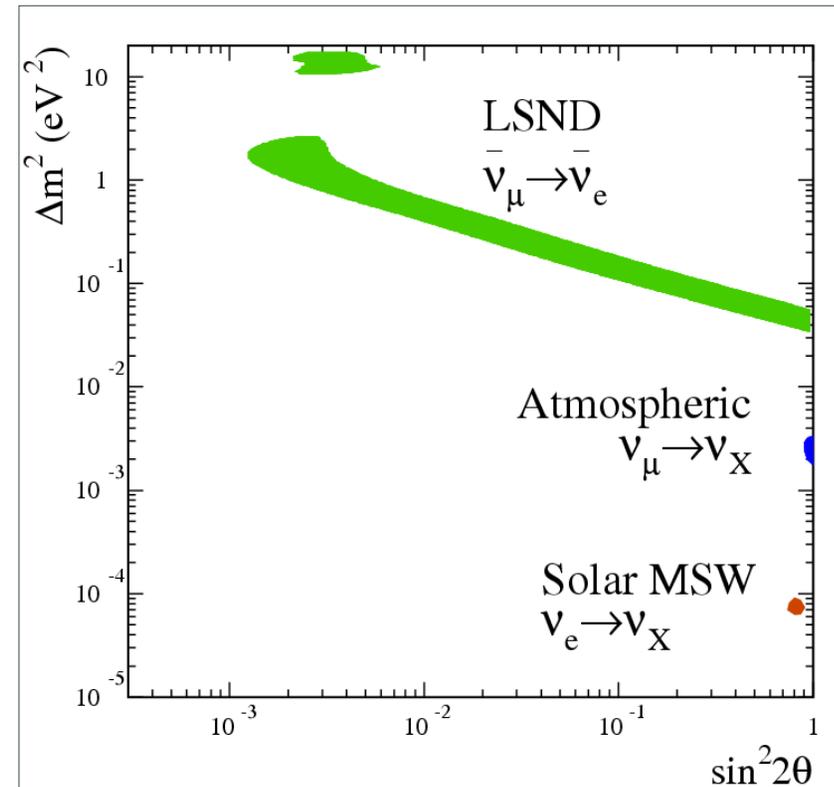
$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\vartheta_{\mu e} \sin^2(1.27\Delta m^2 L/E)$$



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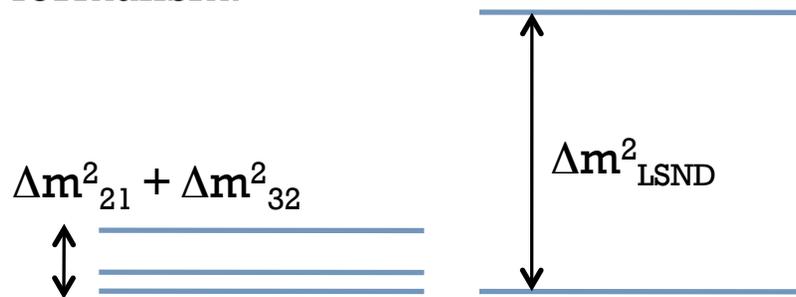
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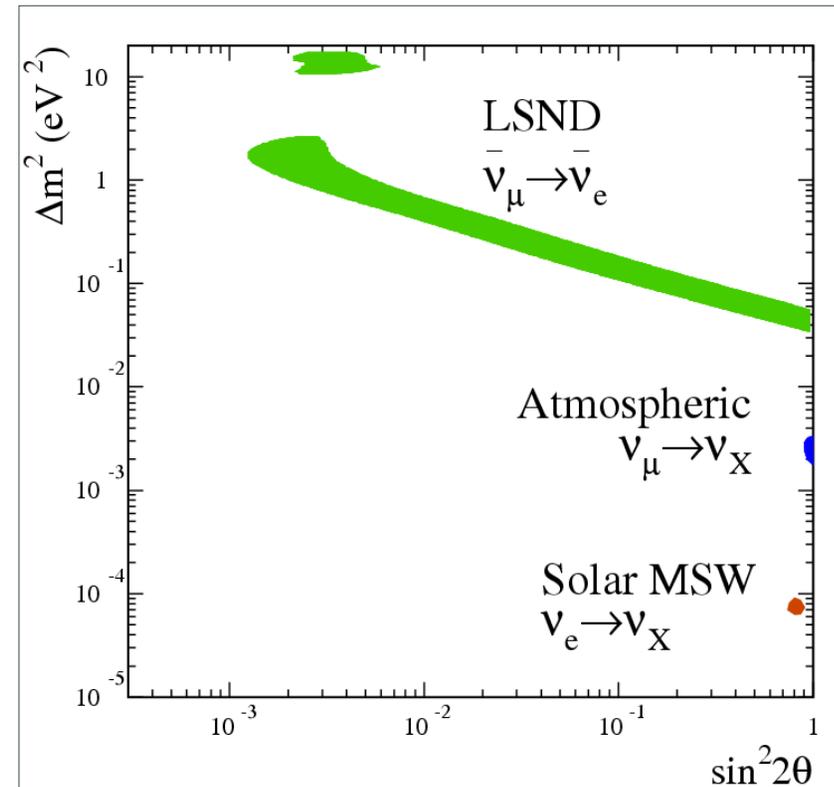
$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\vartheta_{\mu e} \sin^2(1.27\Delta m^2 L/E)$$

In conflict with three-neutrino  
formalism!



$$\Delta m^2_{\text{LSND}} \gg \Delta m^2_{21} + \Delta m^2_{32}$$

**Needs more than 3 neutrinos!**

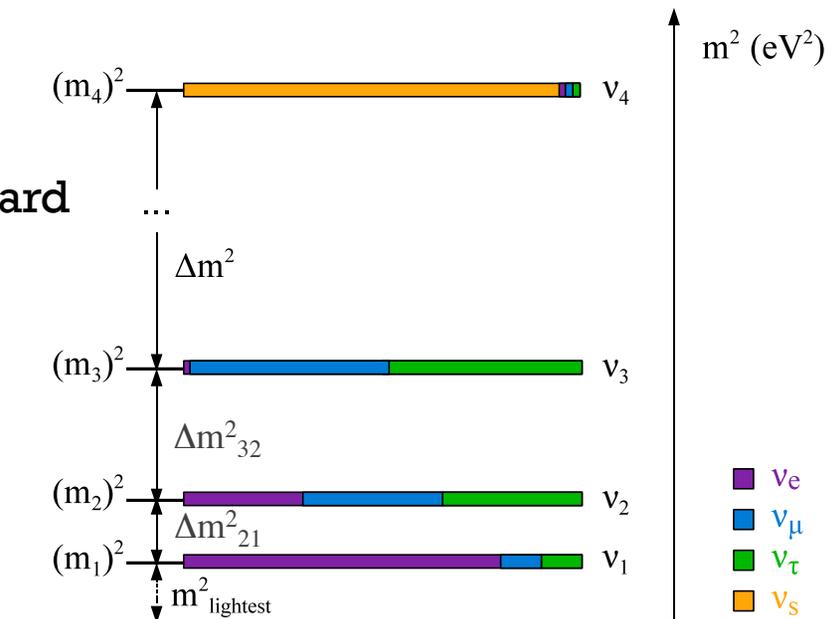


# Possible interpretation: sterile neutrino

Additional neutrino “flavor” (and mass) state which has **no weak interactions** (through the standard W/Z bosons)

Additional mass state is assumed to be produced through mixing with the standard model neutrinos

→ Can affect neutrino oscillations through mixing



# Sterile Neutrino Oscillation Formalism

## Oscillation effects:

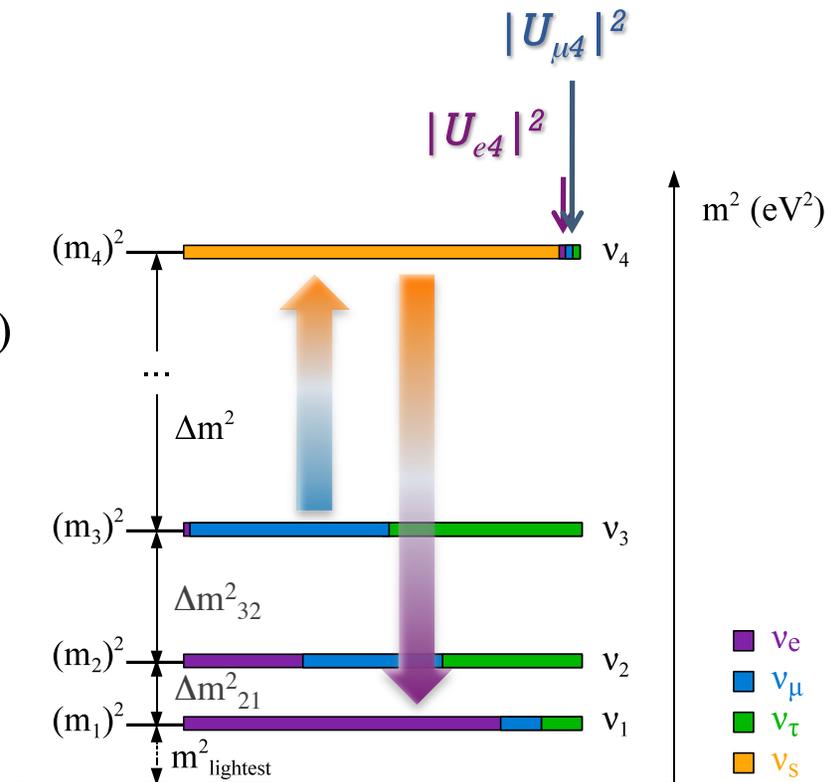
$\nu_\mu \rightarrow \nu_e$  appearance\*:

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\vartheta_{\mu e} \sin^2(1.27\Delta m^2 L/E)$$

$$\downarrow \quad \quad \quad 4|U_{e4}|^2|U_{\mu4}|^2$$

**Explains LSND result  
but needs**

**independent confirmation!**

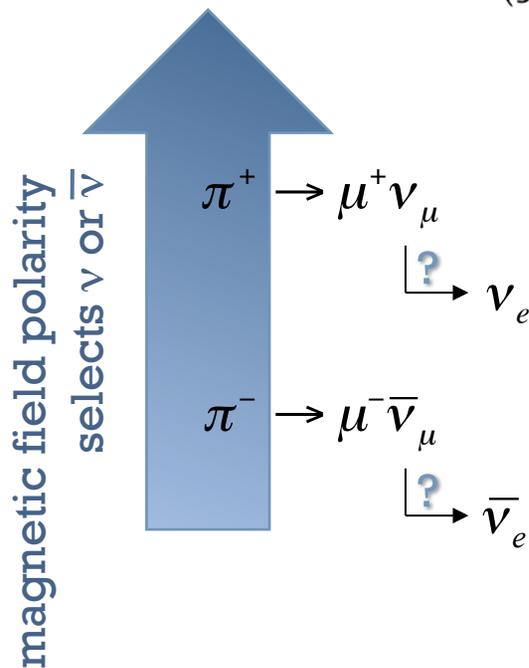
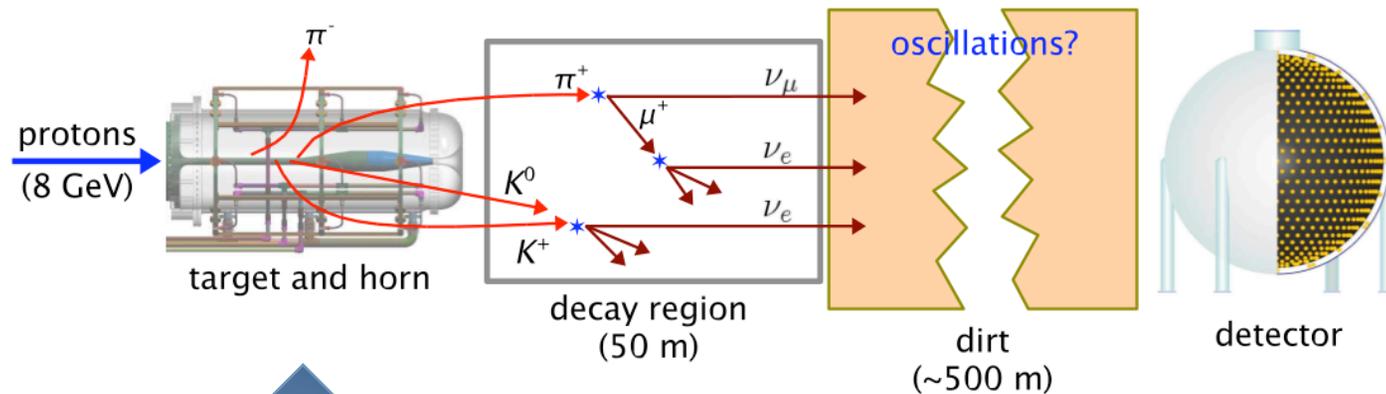


**(3+1)**

\*Approximation:  $m_1, m_2, m_3 \ll m_4 \rightarrow m_1, m_2, m_3 = 0$

# Puzzle piece #2: MiniBooNE

MiniBooNE was proposed to test the LSND result:

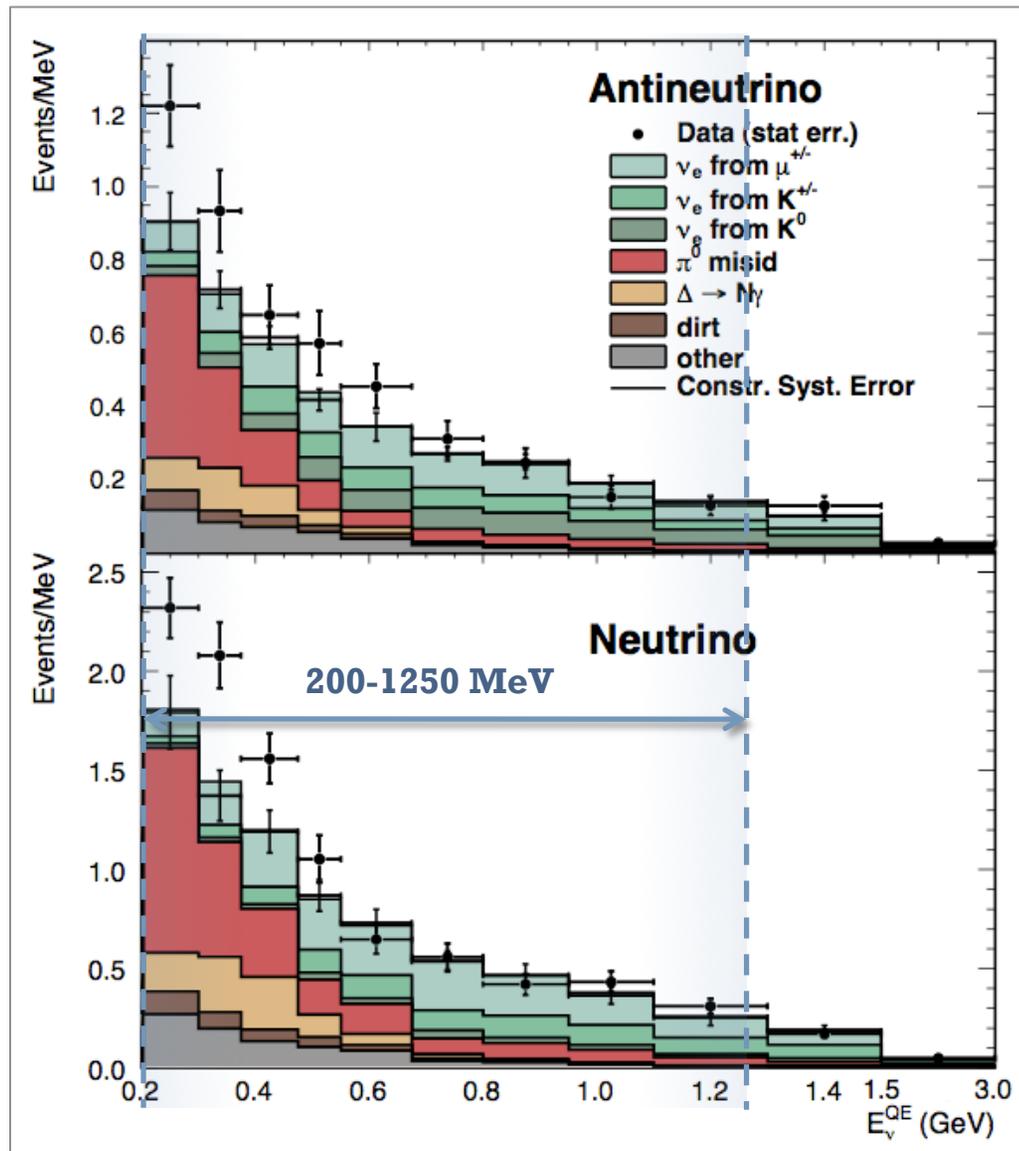


Similar L/E as LSND

but

- Different energy, beam and detector systematics
- Different event signatures and backgrounds (cherenkov detector)

# Puzzle piece #2: MiniBooNE



[arXiv:1303.2588, accepted by Phys. Rev. Lett.; see also:

Phys.Rev.Lett.110.161801,2012

Phys.Rev.Lett.98.231801,2007,

Phys.Rev.Lett.102.101802,2009,

Phys.Rev.Lett.103.111801,2009,

Phys.Rev.Lett.105.181801,2010]

**Oscillation signal region:  
200-1250 MeV**

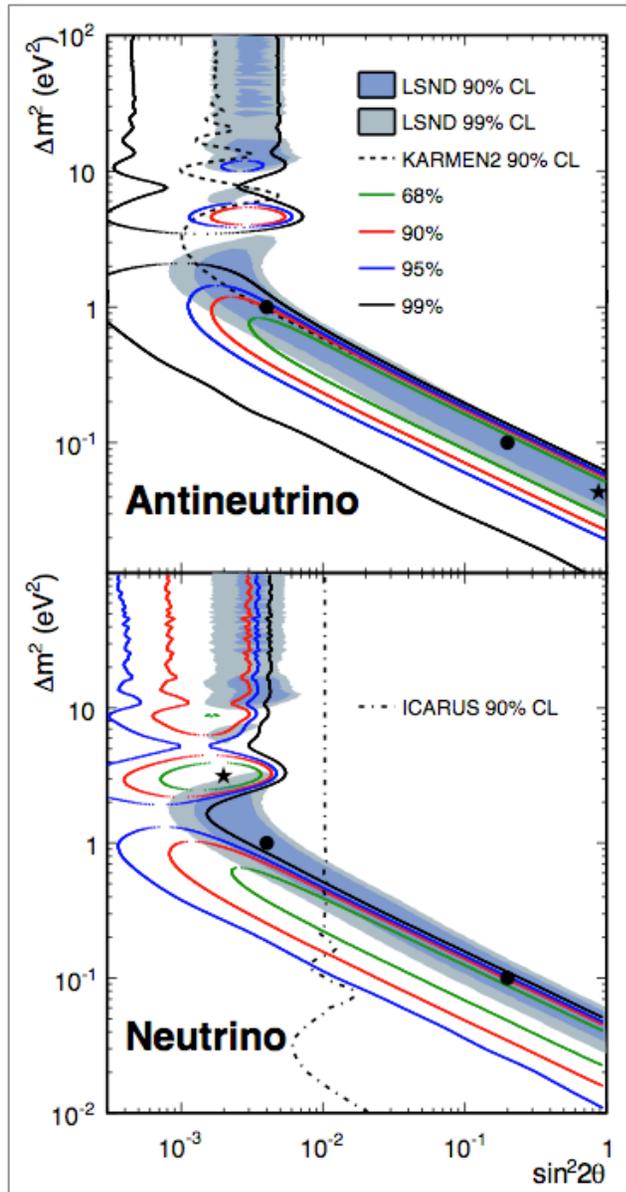
**Antineutrino search:  
2.8 $\sigma$  excess**

Excess of events is at both high and “low energy.”

**Neutrino search:  
3.4 $\sigma$  excess**

Excess of events is at “low energy,”  
 $E < 475$  MeV.

# Puzzle piece #2: MiniBooNE



[arXiv:1303.2588, accepted by Phys. Rev. Let. ;  
see also:

Phys.Rev.Lett.110.161801,2012

Phys.Rev.Lett.98.231801,2007,

Phys.Rev.Lett.102.101802,2009,

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Phys.Rev.Lett.105:181801,2010]

## Antineutrino (3+1) best fit:

$\chi^2$ -probability = 66%

$(\Delta m^2, \sin^2 2\theta) = (0.04 \text{ eV}^2, 0.88)$

Background-only relative to best fit: 0.5%

## Neutrino (3+1) best fit:

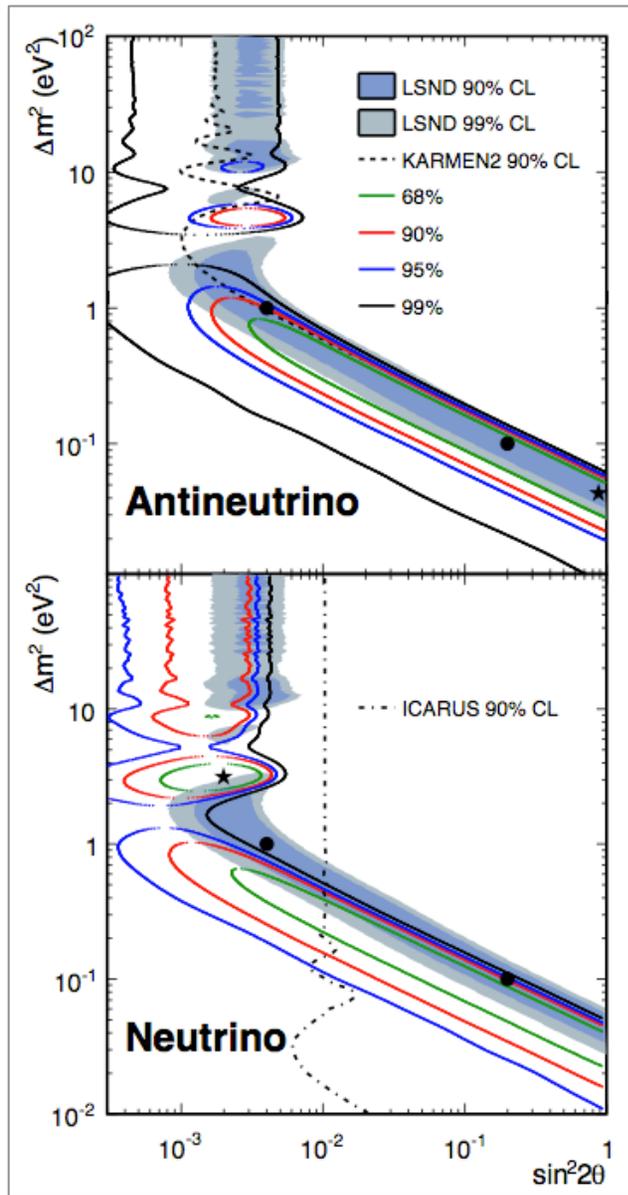
$\chi^2$ -probability = 6.1%

$(\Delta m^2, \sin^2 2\theta) = (3.14 \text{ eV}^2, 0.002)$

Background-only relative to best fit: 2%

**Both are consistent with (3+1) oscillations in general, but MiniBooNE antineutrino allowed parameters are in better agreement with LSND parameters.**

# Puzzle piece #2: MiniBooNE



Barring CP violation,

$$P(\nu_\mu \rightarrow \nu_e) \equiv P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$$

(3+1) approximation  
does not allow for CP violation

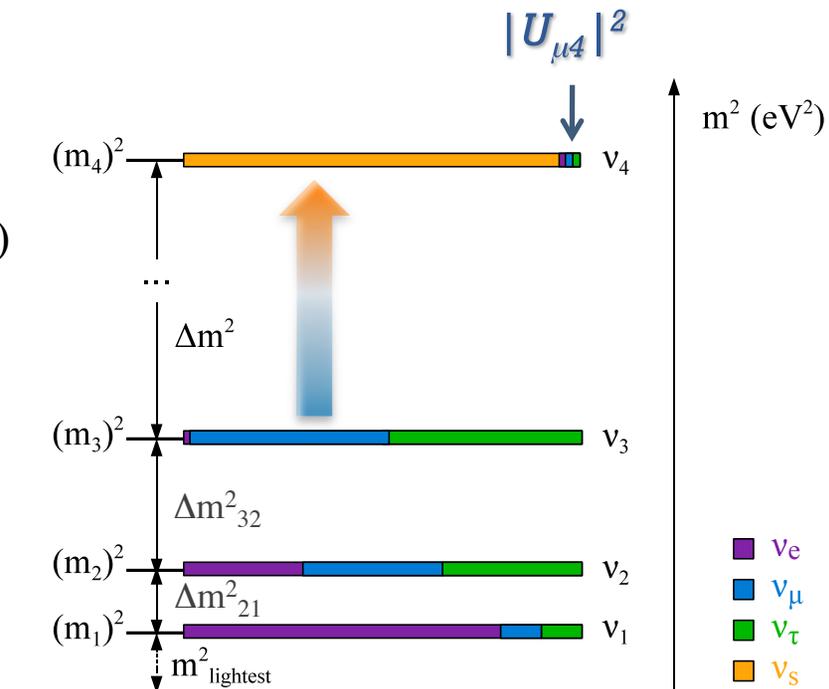
# Sterile Neutrino Oscillation Formalism

$\nu_\mu \rightarrow \nu_e$  appearance implies  $\nu_\mu$  and  $\nu_e$  disappearance!

$\nu_\mu$  disappearance\*:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta_{\mu\mu} \sin^2(1.27\Delta m^2 L/E)$$

$$\hookrightarrow 4|U_{\mu 4}|^2(1 - |U_{\mu 4}|^2)$$



(3+1)

\*Approximation:  $m_1, m_2, m_3 \ll m_4 \rightarrow m_1, m_2, m_3 = 0$

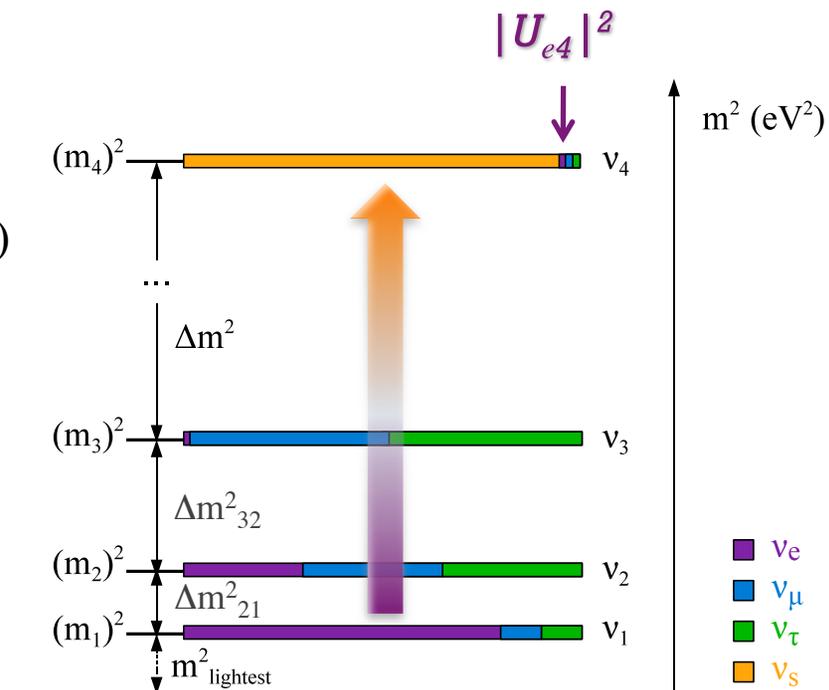
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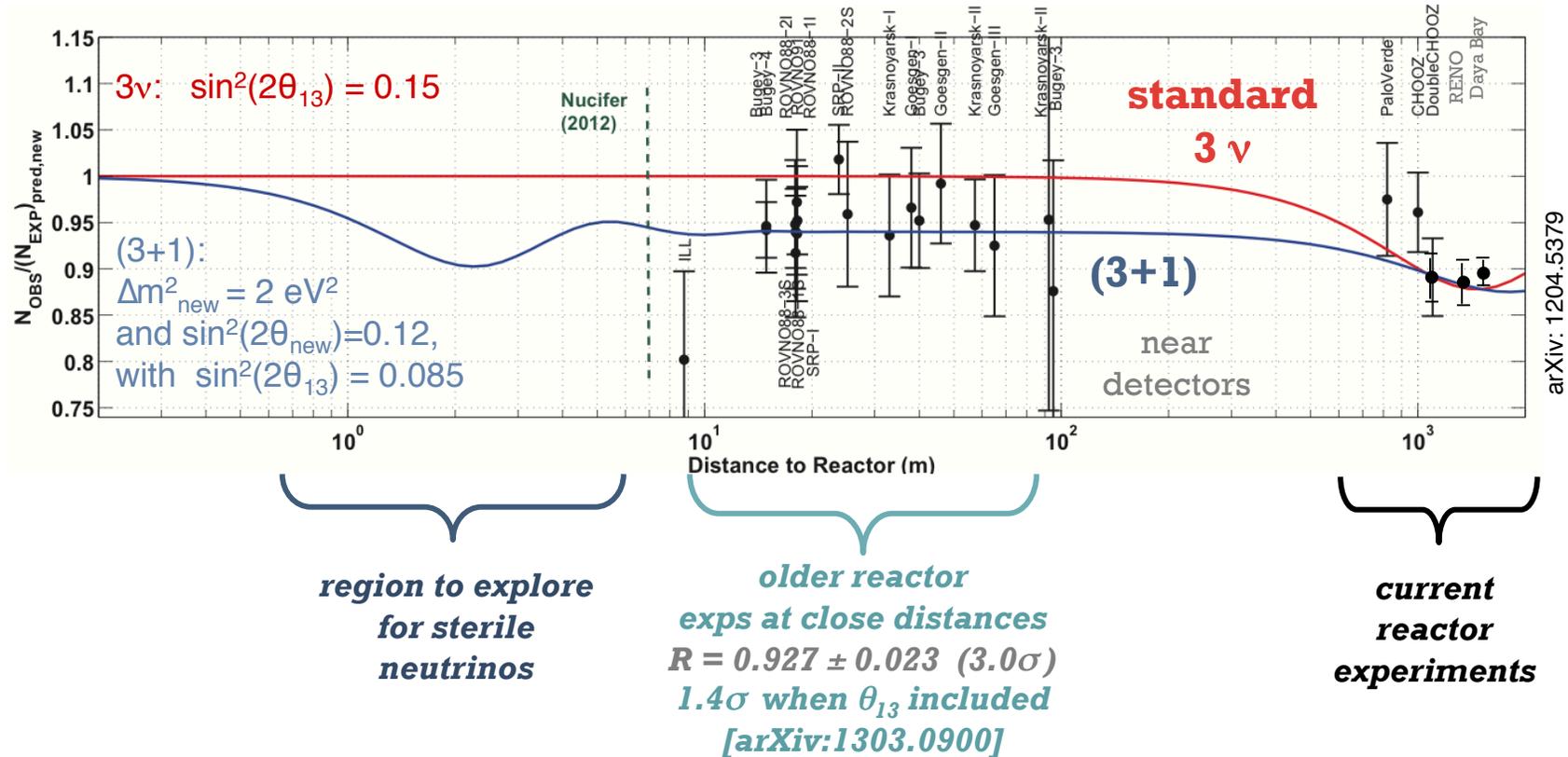


(3+1)

\*Approximation:  $m_1, m_2, m_3 \ll m_4 \rightarrow m_1, m_2, m_3 = 0$

# Puzzle piece #3: Reactor Anomaly

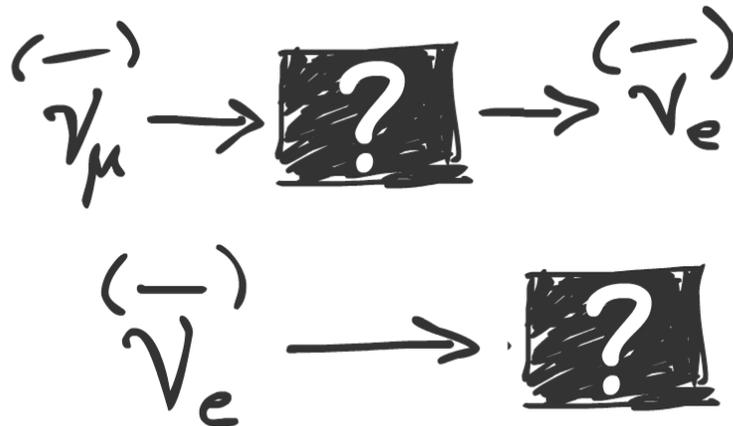
$\bar{\nu}_e \rightarrow \bar{\nu}_s$  disappearance?



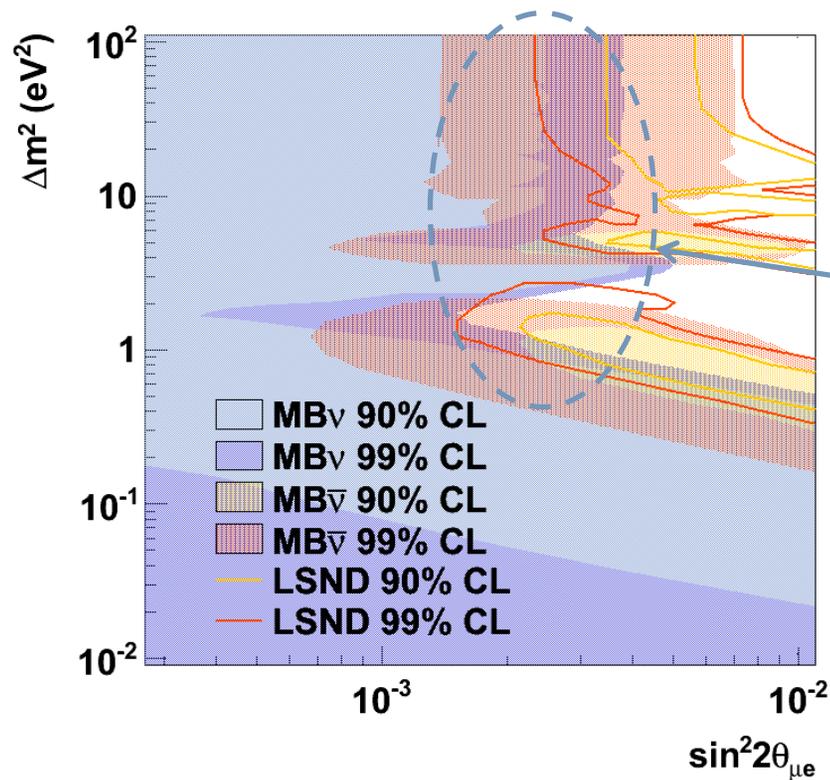
Fewer reactor neutrinos than expected at short baselines

→ A possible interpretation: sterile neutrino osc. with  $\Delta m^2 \sim 1 \text{ eV}^2$  and  $\sin^2 2\theta \sim 0.1$

1. Can all three signatures be explained by (3+1) sterile neutrino hypothesis?



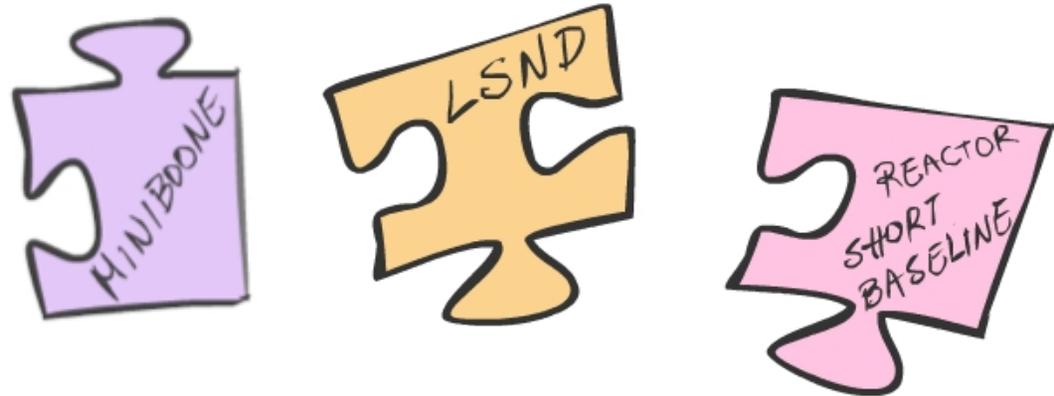
# 1. Can all three signatures be explained by (3+1) sterile neutrino hypothesis?



Reactor short-baseline consistent with these values

**A: Yes!**

1. Can all three signatures be explained by (3+1) sterile neutrino hypothesis?



2. What about information from other experiments sensitive to high- $\Delta m^2$  oscillations?

$(\bar{\nu}_\mu)$  disappearance

CDHS  
CCFR84  
SuperK/K2K (atm)  
MiniBooNE (dis)  
MINOS CC

$(\bar{\nu}_\mu) \rightarrow (\bar{\nu}_e)$  appearance

MiniBooNE  $\nu$   
MiniBooNE  $\bar{\nu}$   
LSND  
KARMEN  
NOMAD  
NuMI-MB

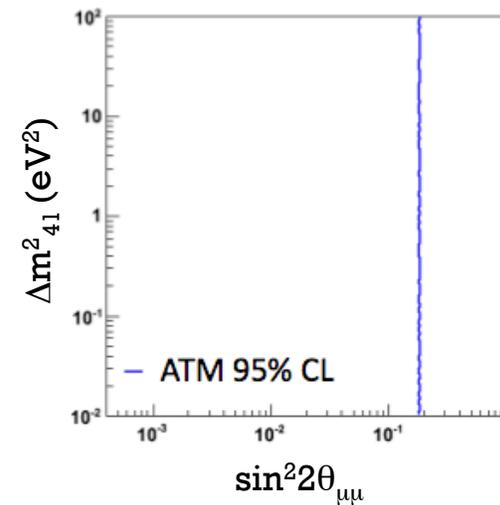
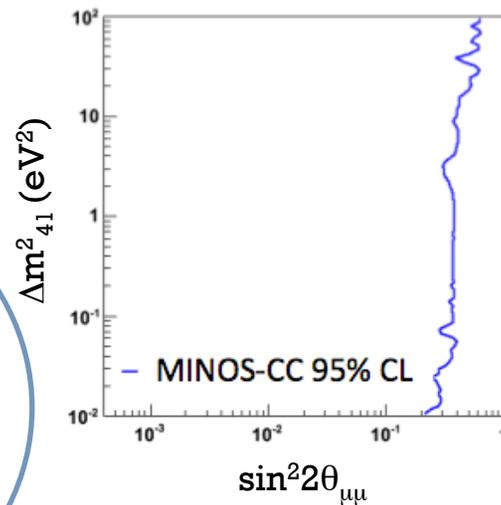
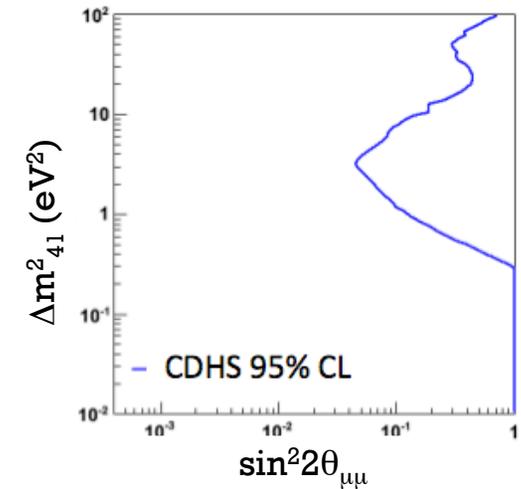
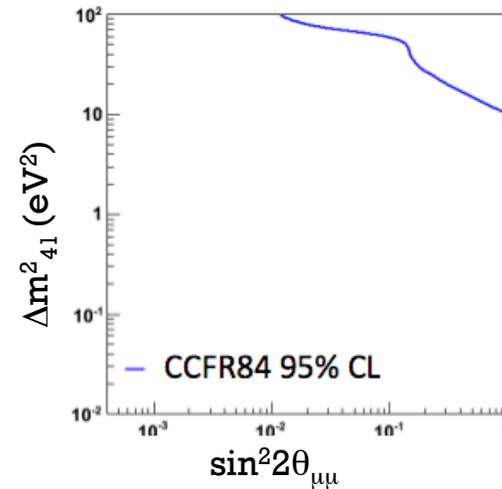
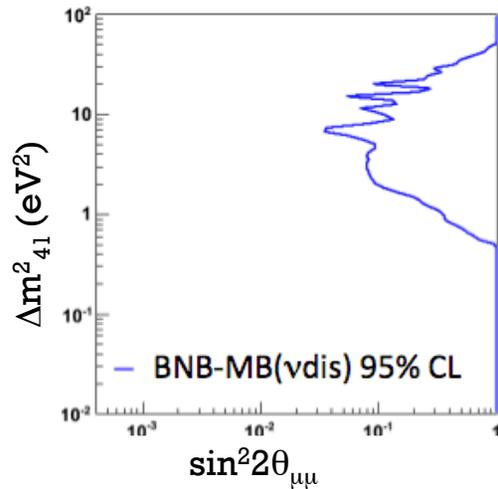
$(\bar{\nu}_e)$  disappearance

Bugey  
KARMEN/LSND (xsec)  
Gallium

[Conrad, Ignarra, GK, Shaevitz, Spitz,  
arXiv:1207.4765, accepted by Advances in HEP;  
see also:

GK et al, Phys.Rev. D80 (2009) 073001,  
GK et al, Phys.Rev. D75 (2007) 013011]

# Other experimental constraints

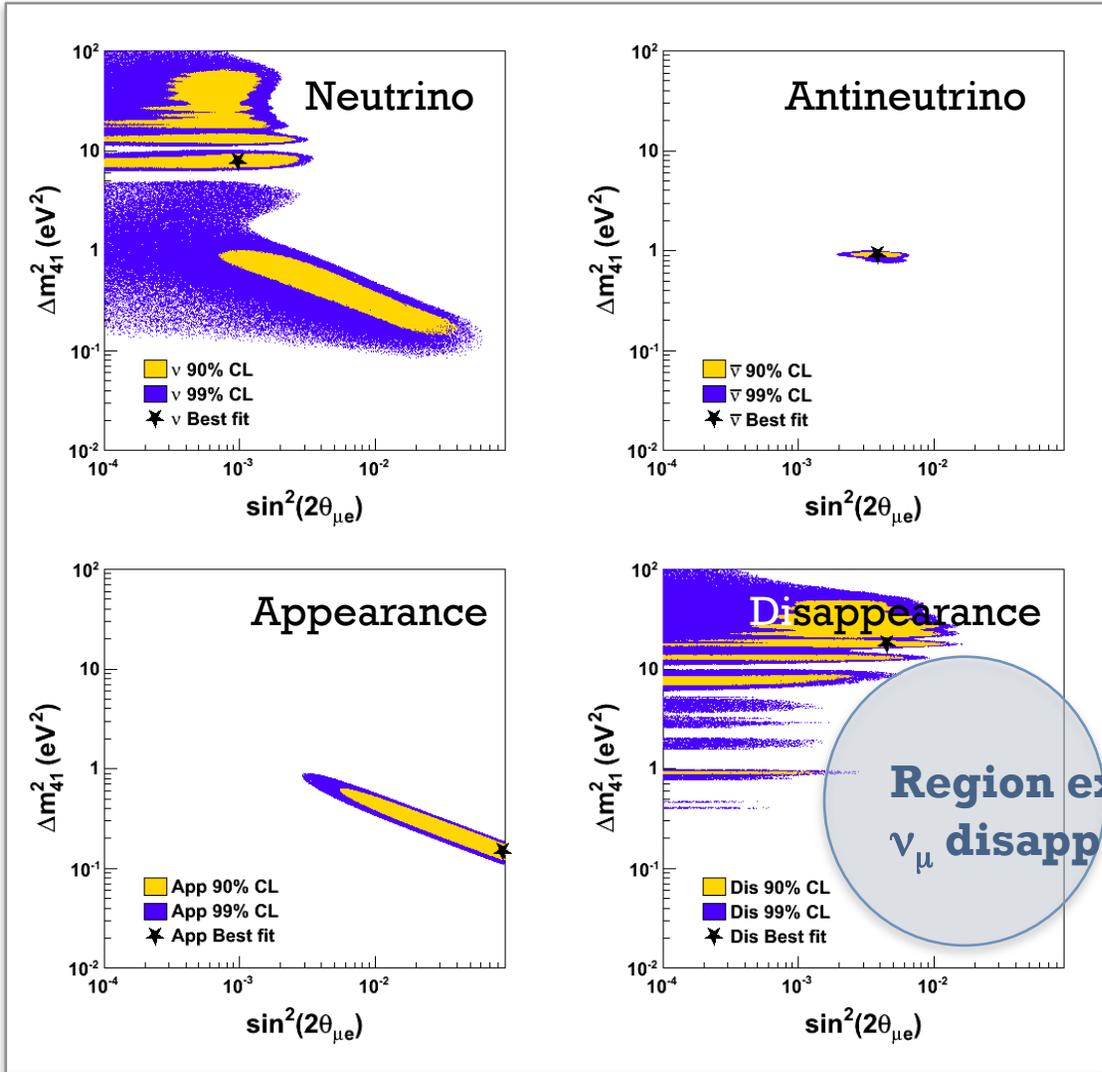


$\nu_{\mu}$  disappearance

CDHS  
CCFR84  
SuperK/K2K (atm)  
MiniBooNE (dis)  
MINOS CC

# (3+1) Global Fits to Sterile Neutrino Oscillations

## Incompatibilities!



Compatibility ( $\nu, \bar{\nu}$ ) = 0.14%

Compatibility (app, dis) = 0.013%

Region excluded from  $\nu_{\mu}$  disappearance experiments

[Conrad, Ignarra, GK, Shaevitz, Spitz, arXiv:1207.4765, accepted by Advances in HEP]

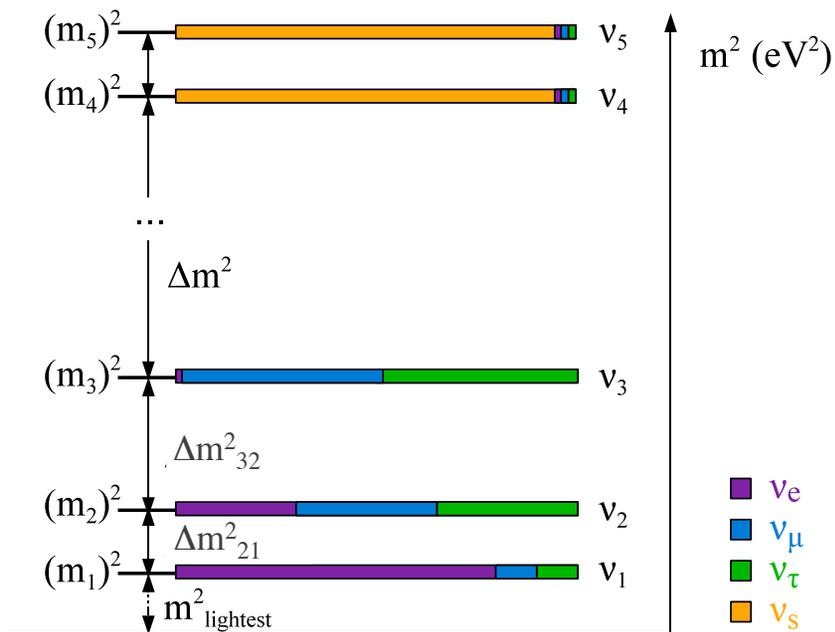
(3+1) is not enough!

Theoretical developments attempting to address inconsistencies:

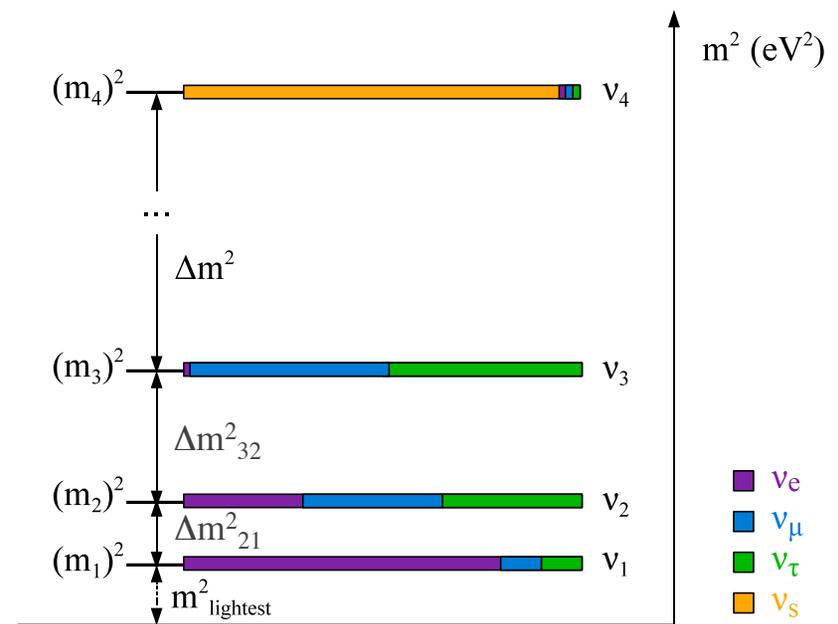
- Fact #1:  $\nu$  vs  $\bar{\nu}$  differences  
Extended sterile neutrino models with CP violation?
- Fact #2: appearance vs disappearance differences  
“Non-standard” oscillations?

# Extended models: (1) CP violation

Can have more than one new state...



(3+2)



(3+1)

# Extended models: (1) CP violation

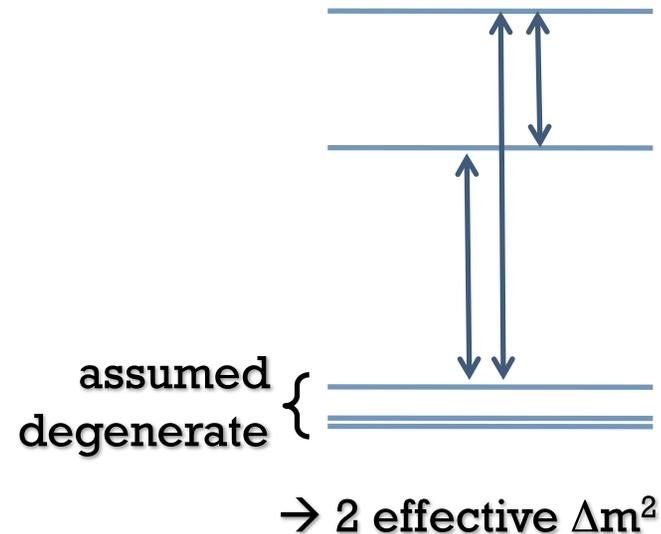
Disappearance:

$$P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - 4[(1 - |U_{\alpha 4}|^2 - |U_{\alpha 5}|^2) \cdot (|U_{\alpha 4}|^2 \sin^2 x_{41} + |U_{\alpha 5}|^2 \sin^2 x_{51}) + |U_{\alpha 4}|^2 |U_{\alpha 5}|^2 \sin^2 x_{54}]$$

Appearance:

$$P(\nu_\alpha \rightarrow \nu_{\beta \neq \alpha}) = 4|U_{\alpha 4}|^2 |U_{\beta 4}|^2 \sin^2 x_{41} + 4|U_{\alpha 5}|^2 |U_{\beta 5}|^2 \sin^2 x_{51} + 8|U_{\alpha 5}| |U_{\beta 5}| |U_{\alpha 4}| |U_{\beta 4}| \sin x_{41} \sin x_{51} \cos(x_{54} - \phi_{45})$$

$x_{ji} \equiv 1.27 \Delta m_{ji}^2 L/E$  CPV phase

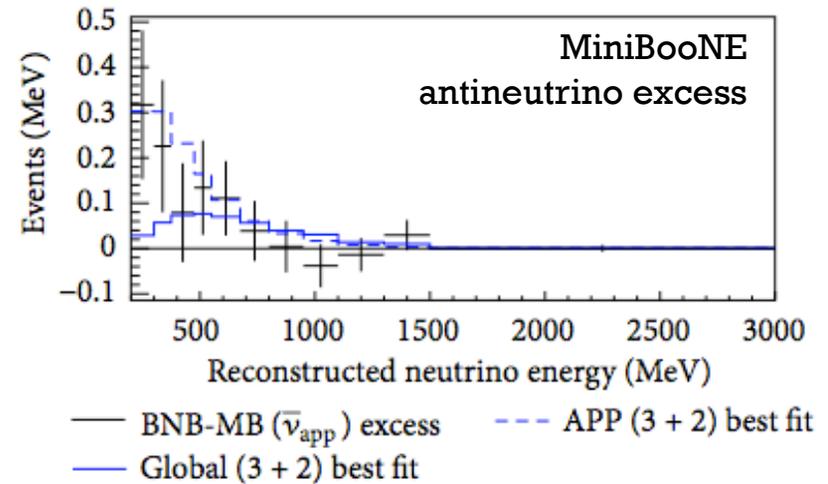
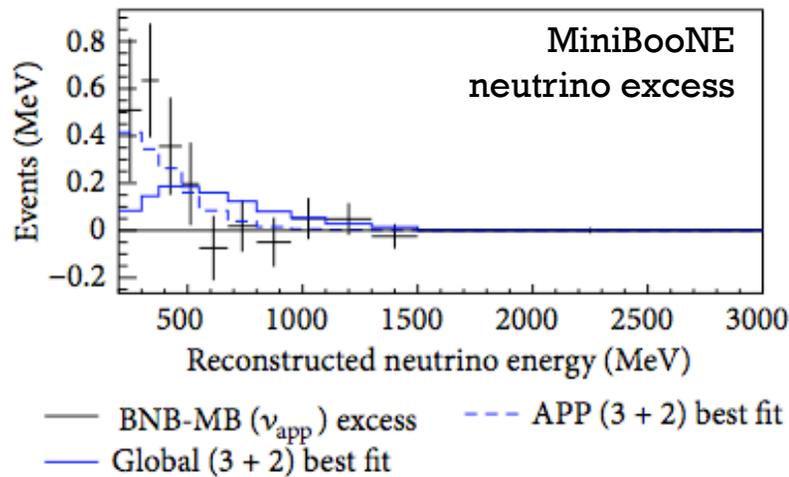


**(3+2)** is attractive because of  
**CP violation**

# Extended models: (1) CP violation

[Conrad, Ignarra, GK, Shaevitz, Spitz,  
arXiv:1207.4765, accepted by Advances in HEP]

(3+2) global best fit



**(3+2) with CP violation cannot explain  
MiniBooNE low E excess, unless  
we throw out disappearance  
constraints!**

# Global Fits: Caveats and Limitations

- Appearance searches assume no disappearance
  - This is an incorrect assumption, given best fit parameters extracted in global fits
  - This may resolve some tension seen in the MiniBooNE appearance data sets, if one allows for  $\nu_e$  background disappearance
- Need a more advanced statistical and systematic treatment of data sets
  - Compatibility measure needs to be verified with fake data and frequentist studies
  - Need better treatment of systematic correlations between data sets.

This is a challenging step, but necessary for meaningful quantitative statements on these models

## Theoretical developments attempting to address inconsistencies:

- ~~Fact #1:  $\nu$  vs  $\bar{\nu}$  differences~~  
~~Extended sterile neutrino models with CP violation?~~
- Fact #2: appearance vs disappearance differences  
“Non-standard” oscillations?

*Does not explain  
MINIBOONE low E excess*

## Theoretical developments attempting to address inconsistencies:

- ~~Fact #1:  $\nu$  vs  $\bar{\nu}$  differences~~  
~~Extended sterile neutrino models with CP violation?~~
- ~~Fact #2: appearance vs disappearance differences~~  
~~“Non-standard” oscillations?~~

*Does not explain  
MINIBOONE low E excess*

*MINIBOONE low E excess??  
difficult to interpret...*

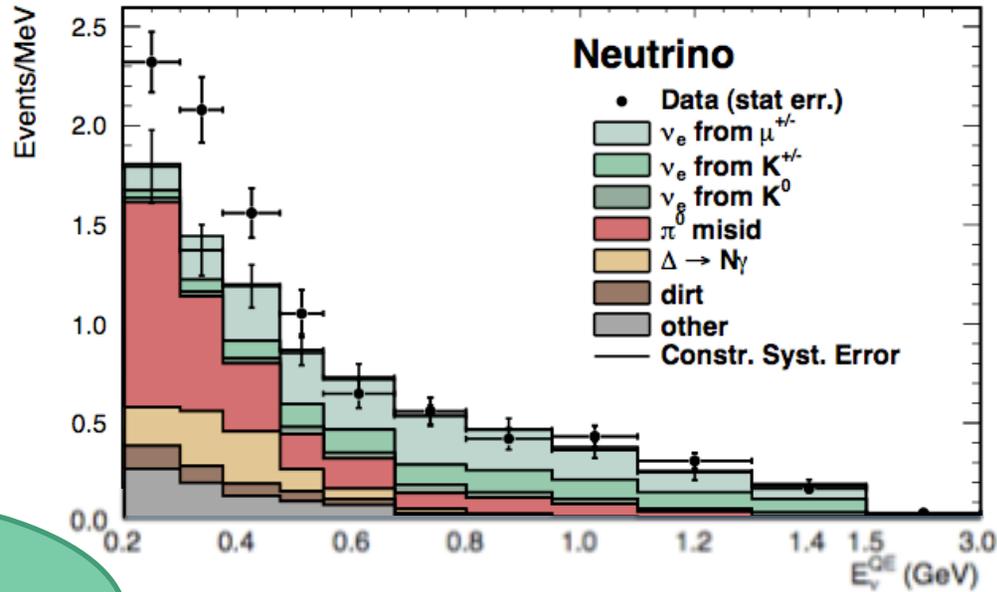
### Other theoretical interpretations:

- CPT violation
- Heavy (sterile) neutrino decay
- Extra dimensions
- New interactions
- Altered neutrino dispersion relations

None of these  
work that great



### Low E excess: Are we missing something?



Unaccounted  $\nu_e/\nu_\mu$  disappearance?

Energy reconstruction?  
Cross-section/  
nuclear effects?

Electron-like misestimated or new background?

Single-photon mis-estimated or new background?

# Big Questions in Neutrino Physics

What is the value of  $\delta_{CP}$ ?

Is the neutrino mass spectrum normal, or inverted?

What are the absolute neutrino masses?

Are neutrinos dirac or majorana fields?

**Are there additional, “sterile” neutrino states?**

Do we understand exclusive and inclusive neutrino cross sections on nuclear targets?

Fundamental  
questions

Pressing  
experimental  
questions

# Big Questions in Neutrino Physics

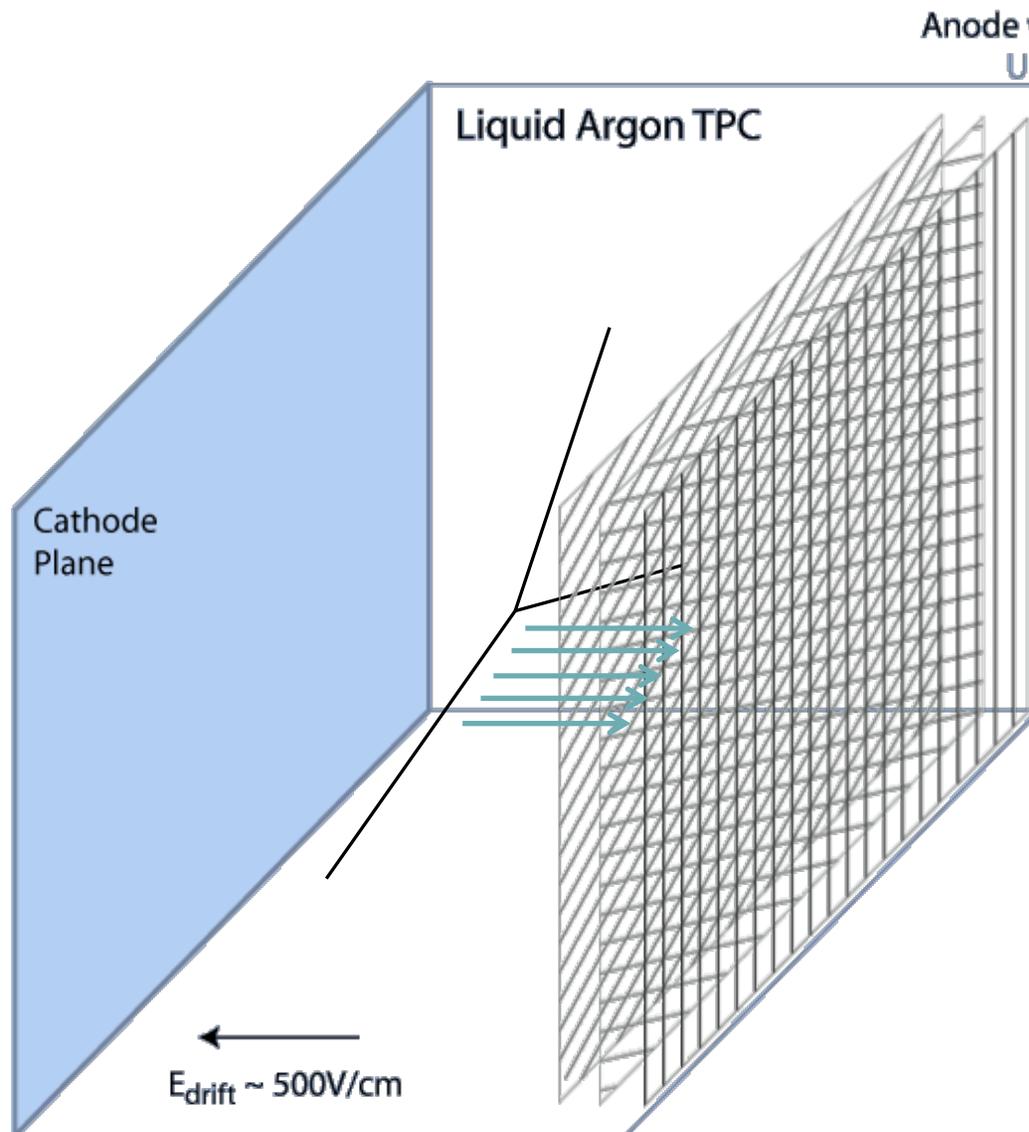
Directly addressed  
by LArTPC oscillation  
experiments?

- ✓ What is the value of  $\delta_{CP}$ ?
- ✓ Is the neutrino mass spectrum normal, or inverted?  
What are the absolute neutrino masses?  
Are neutrinos dirac or majorana fields?
- ✓ **Are there additional, “sterile” neutrino states?**
- ✓ Do we understand exclusive and inclusive neutrino cross sections on nuclear targets?

Fundamental  
questions

Pressing  
experimental  
questions

# LArTPC detector concept



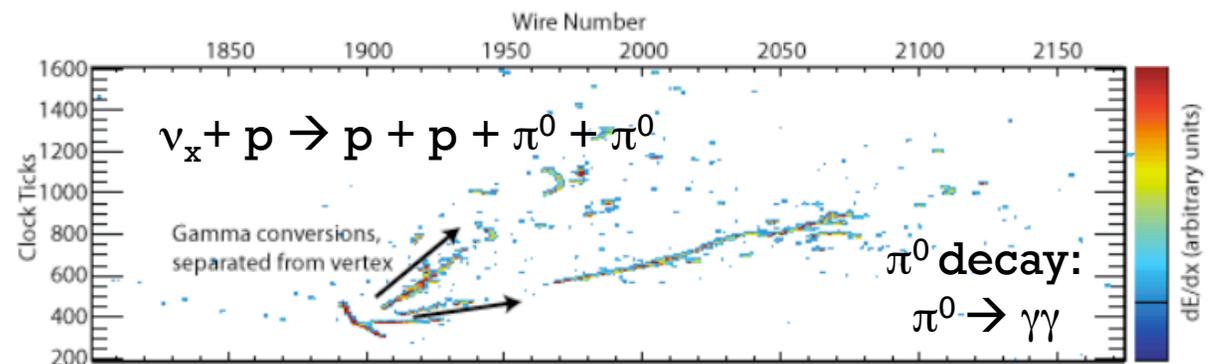
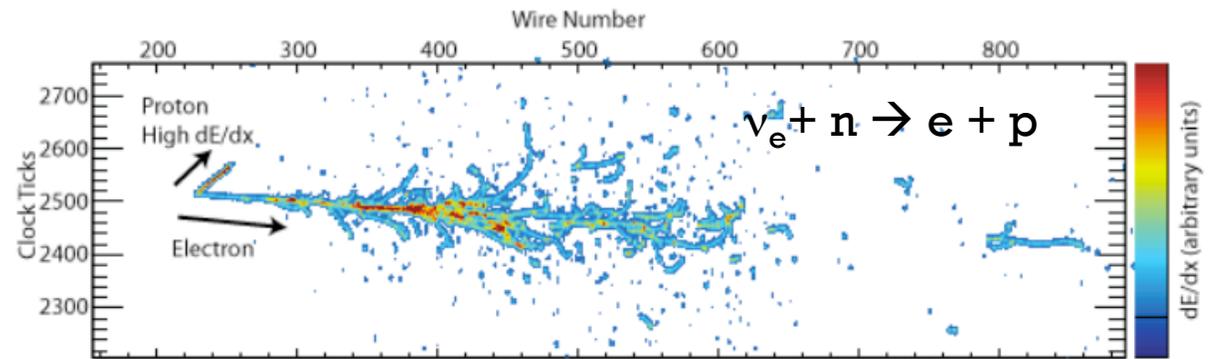
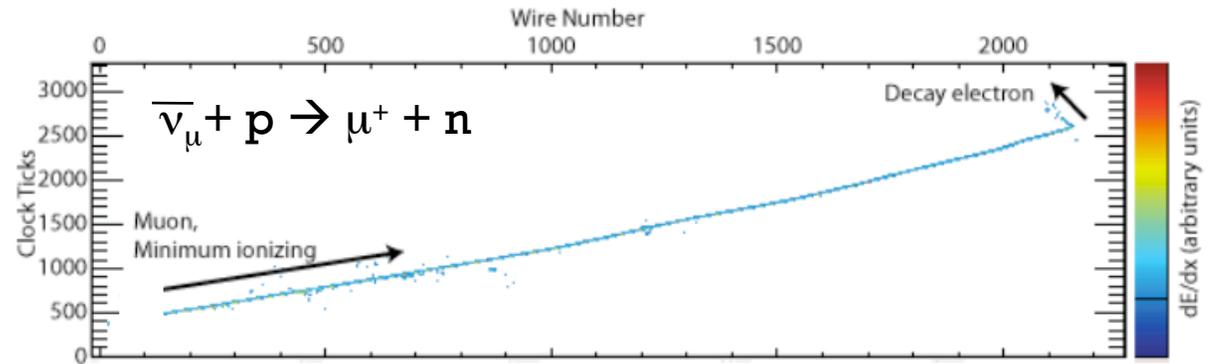
Charged particle tracks ionize argon atoms; **Ionization charge** drifts to **finely segmented charge collection planes** over  $\sim 1$ -few ms.



Scintillation light ( $\sim$ few ns) is detected by photo-sensitive detectors for event  $t_0$ , drift coordinate and triggering

# Exquisite event topology!

High event selection efficiency and excellent background rejection!



# e/ $\gamma$ differentiation

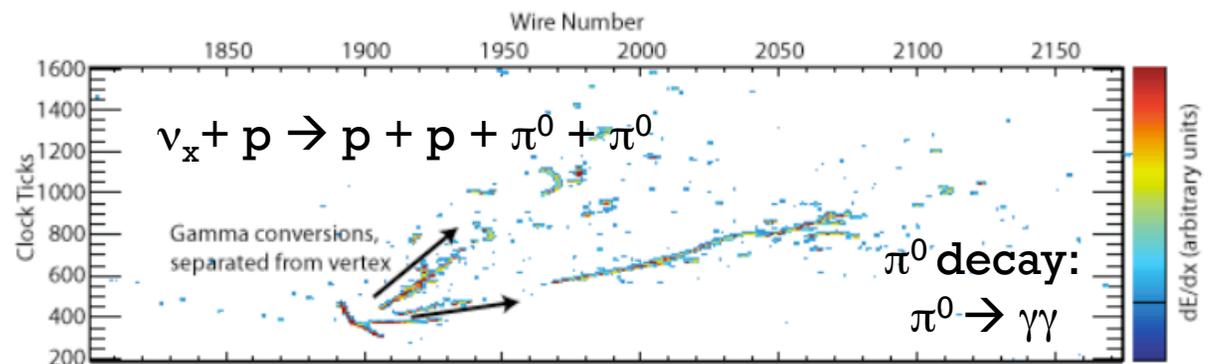
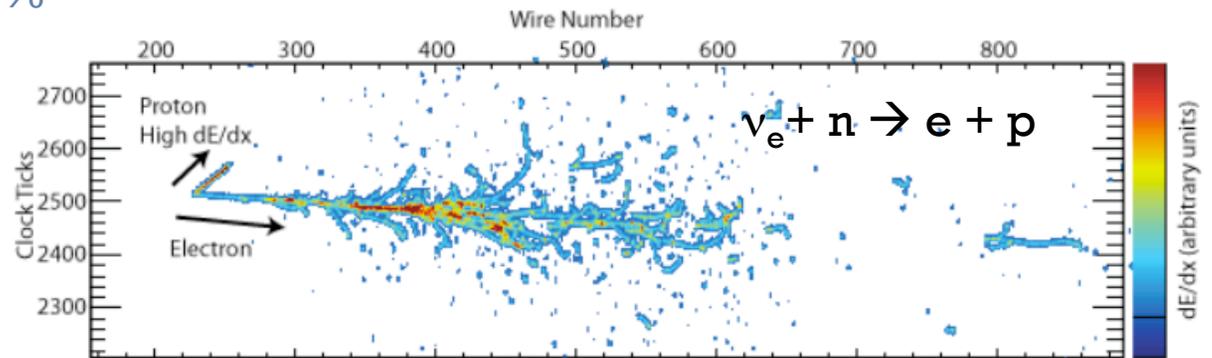
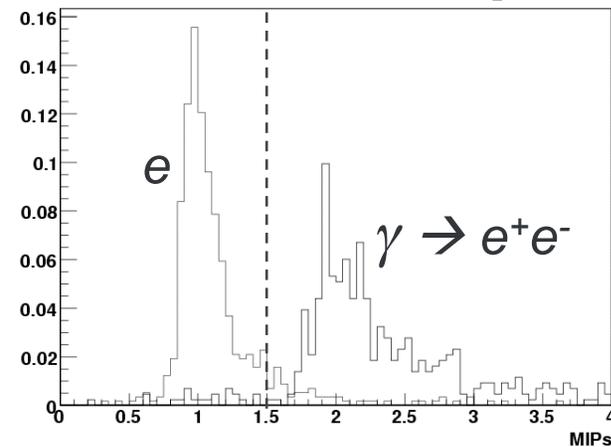
$\gamma$ 's rejected on the basis of

1. detached shower vertex
2. larger dE/dx at the beginning of shower

Typical e/ $\gamma$  separation:  $\sim 90\%$

→ Ideal technology for  $\nu_e$  measurements!

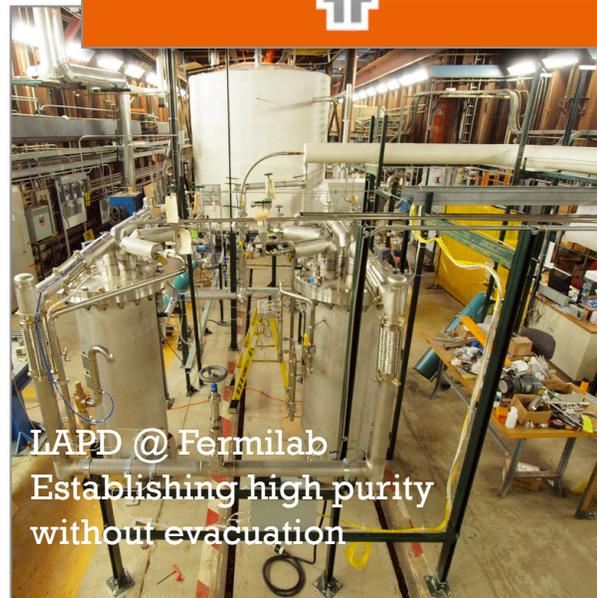
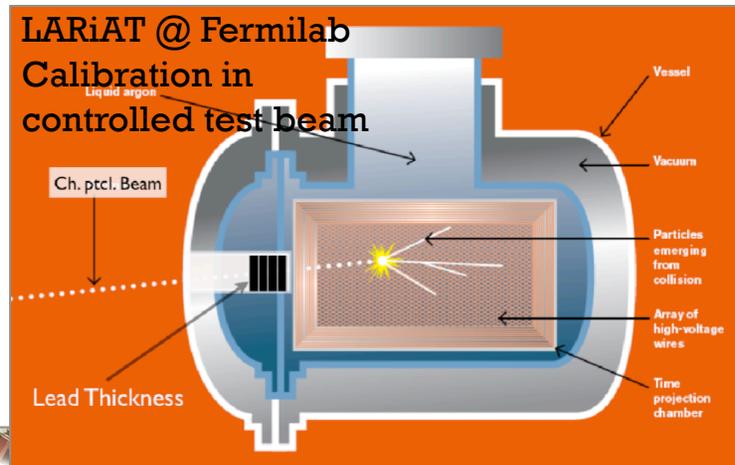
Energy loss in first 24mm of track:  
250 MeV electron vs. 250 MeV photon



# Scalability is challenging!

[Being addressed by ongoing and planned R&D projects]

- Large cryogenic system
- Long drift distances
  - Requires ultra high purity and evacuation is impractical
  - Implies high voltage on cathode
- Large number of readout channels with high data volume/channel (data storage, data processing, ...)
- Cold electronics
- Reconstruction tools



# LArTPC's: Test Facilities & Experiments

## United States

Materials Test Stand

ArgoNeuT

LAPD

MicroBooNE

LArI

LARiAT

CAPTAIN

GLADE

RADAR

LBNE

## Europe

50-liter @ CERN

10m<sup>3</sup>

ICARUS

LArTPC in B-field

LANDD @ CERN

ArgonTube @ Bern

UV Laser

2-LAr @ CERN-SPS

GLACIER/LAGUNA

## Japan

Test-Beam (T32) at J-PARC  
100 kton @ Okinoshima island

# MicroBooNE

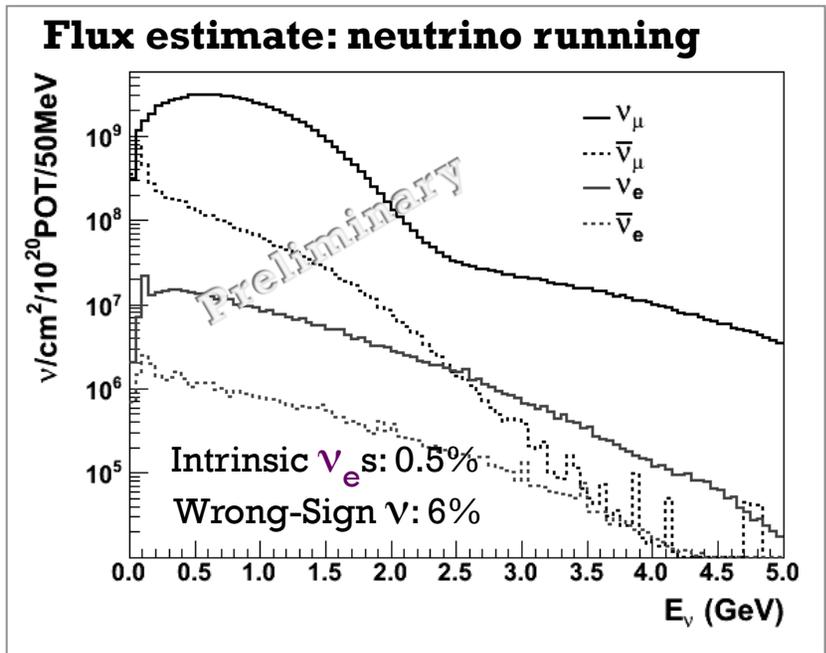
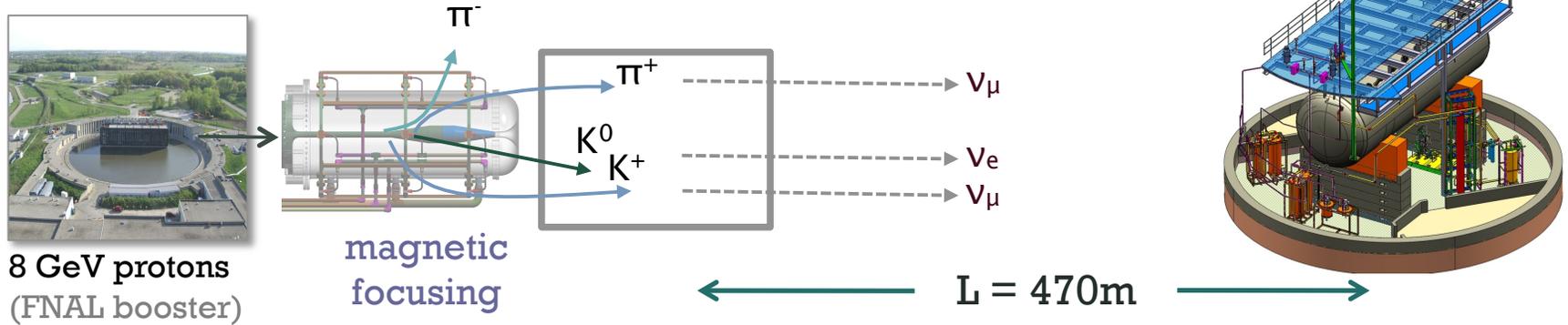
First large-scale LArTPC in the US!



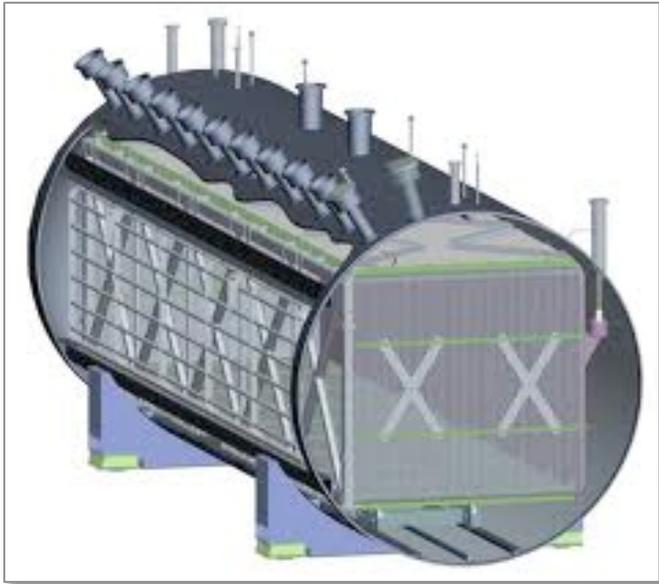
MicroBooNE cryostat at Fermilab on March 8, 2013

# MicroBooNE

Located in the **Fermilab Booster Neutrino Beamline**:

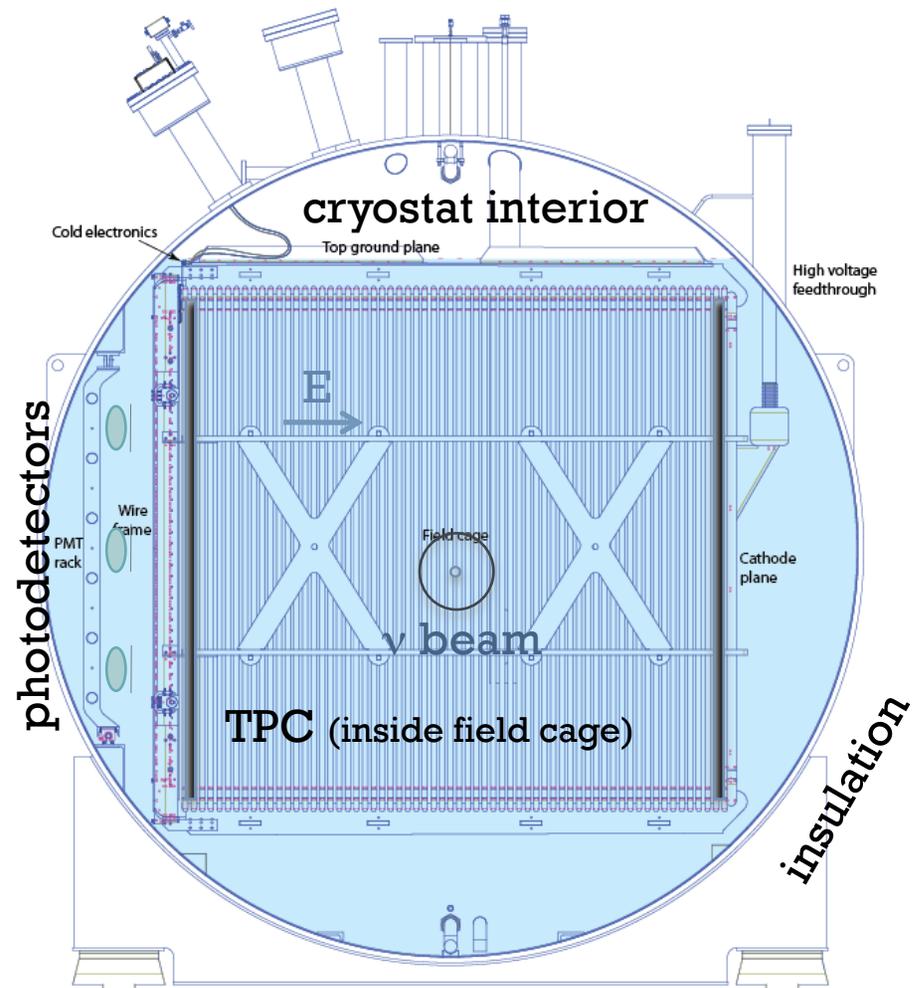


Current run plan (approved):  
Neutrino mode running,  $6.6 \times 10^{20}$  POT



- Detector parameters:
  - 2.5 m x 2.3 m x 10.2 m TPC
  - 2.5 m drift length
  - 170 (60) tons total (fiducial) mass
  - 3 wire planes,  $0, \pm 60^\circ$  from vertical, 3 mm wire separation
  - 8256 wires
  - 32 PMT's for  $t_0$ , drift coordinate, and triggering for empty beam spill rejection

Cross section of detector:



# MicroBooNE

## A lot of data!

**Each event: 160 MB**

8256 wires

read over 4.8 ms

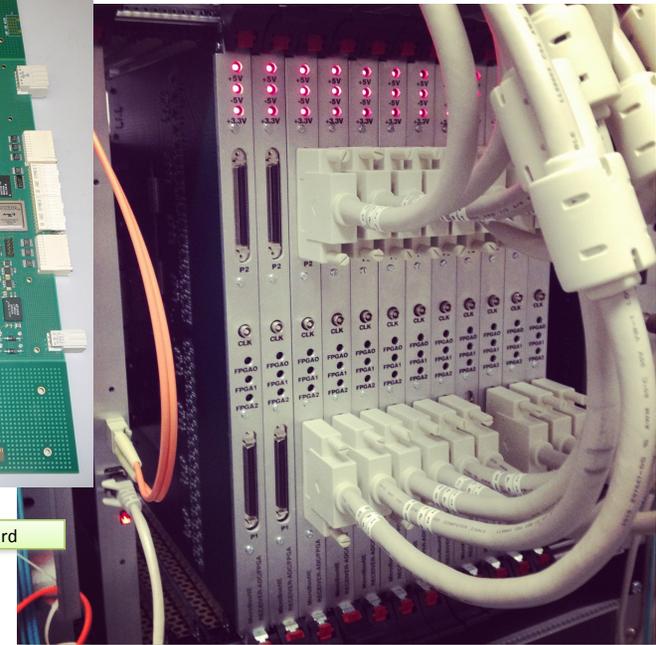
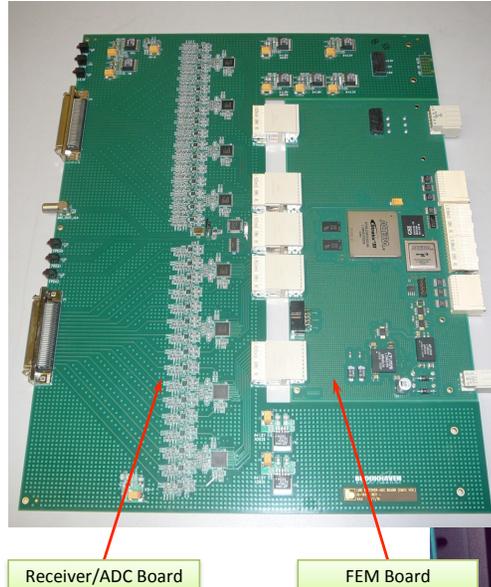
digitized at 2 MHz

12-bit ADC (16-bit packets)

Event rate of  $\sim 0.1$ -10 Hz

→ need compression:

Huffman (lossless) compression provides up to x15 reduction



MicroBooNE readout electronics  
developed at Nevis Labs



**100% live for SuperNova neutrino search: >30 GB/s !**

Solution:

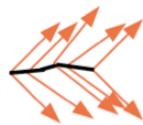
Implement zero suppression, Huffman compression,  
and only retain  $\sim$ few hrs of data at any time

## Primary physics goal

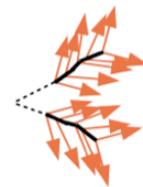
- Investigate the nature of the MiniBooNE low energy excess
- Is the excess due to  $e$  or  $\gamma$ ?

Single  $e$  and single  $\gamma$  are indistinguishable in a cherenkov detector...

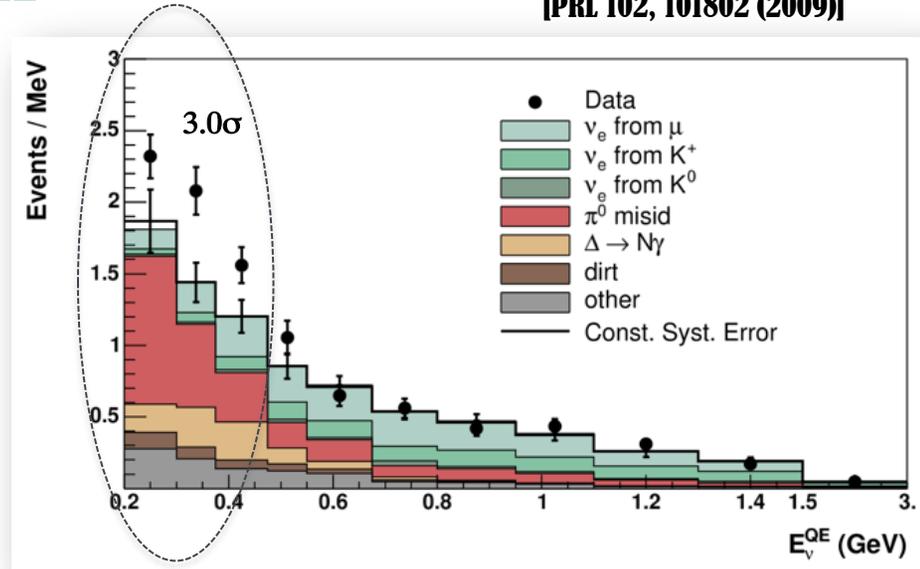
electron:  
short track,  
multiple scattering,  
bremsstrahlung



photon(s):  
photoconversion  
→ electron-like track(s)

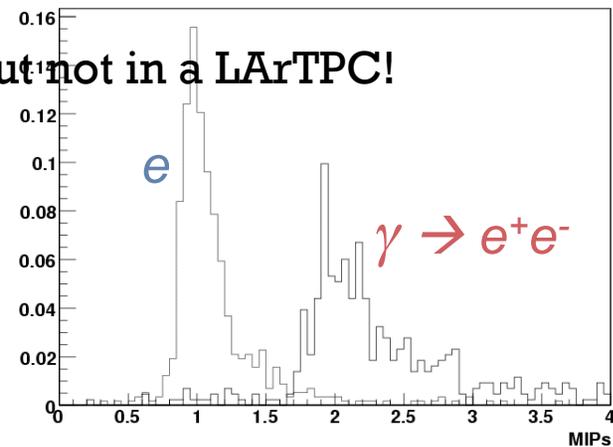


MiniBooNE unexplained “low energy excess”  
[PRL 102, 101802 (2009)]



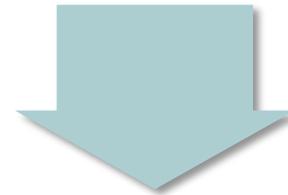
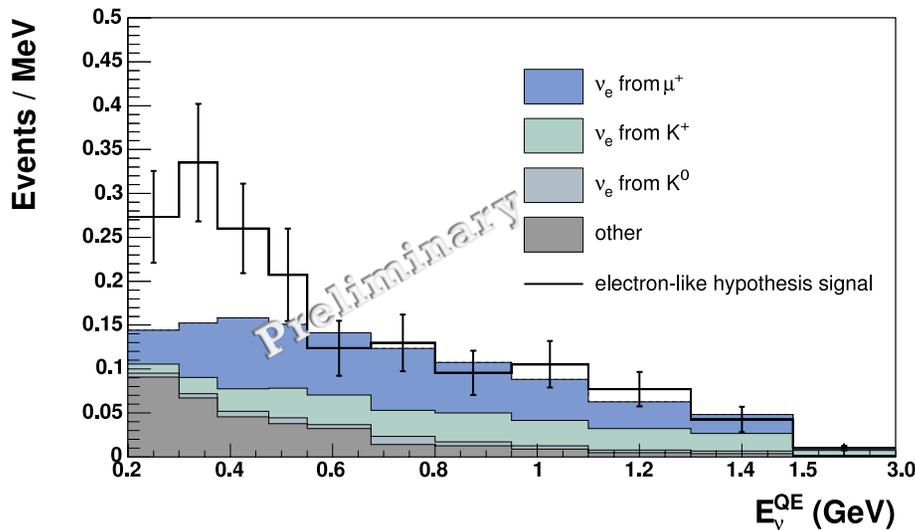
Energy loss in first 24mm of track:  
250 MeV electron vs. 250 MeV photon

...but not in a LArTPC!



## Primary physics goal

What MicroBooNE expects to see if excess is due to **single e**

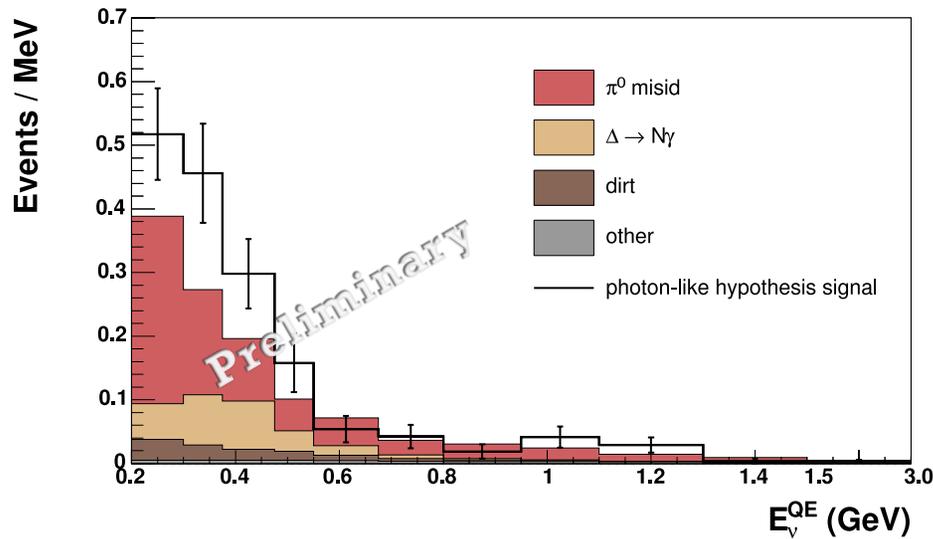


Possible explanation:  
 $\nu_\mu \rightarrow \nu_e$  nonstandard  
oscillations  
(sterile neutrinos, extra  
dimensions, NSI,...)

About **37 excess events** above a background of 45 events  
 $\rightarrow$  **5.7 $\sigma$  statistical significance**

## Primary physics goal

What MicroBooNE expects to see if excess is due to **single  $\gamma$**

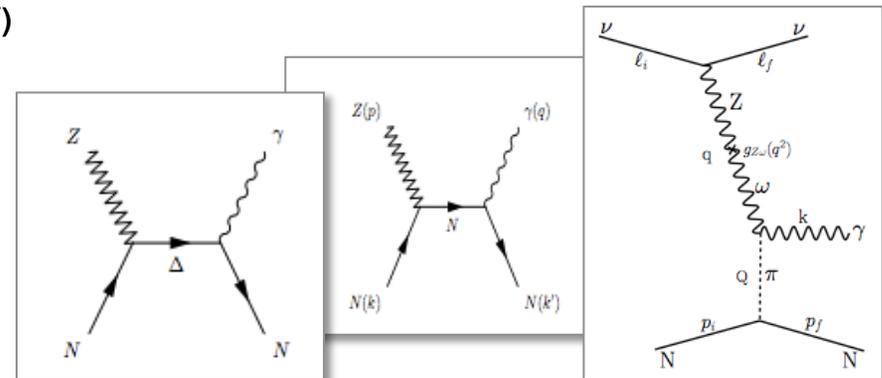


Possible explanation:  
background  $\gamma$  or  $\pi^0$  or  
“new” single photon  
production

e.g.

- R. Hill arXiv: 0905.0291
- Jenkins et al arXiv:0906.0984
- Serot et al arXiv: 1011.5913

About **37 excess events** above a  
background of 79 events  
 $\rightarrow$   **$4.1\sigma$  statistical significance**

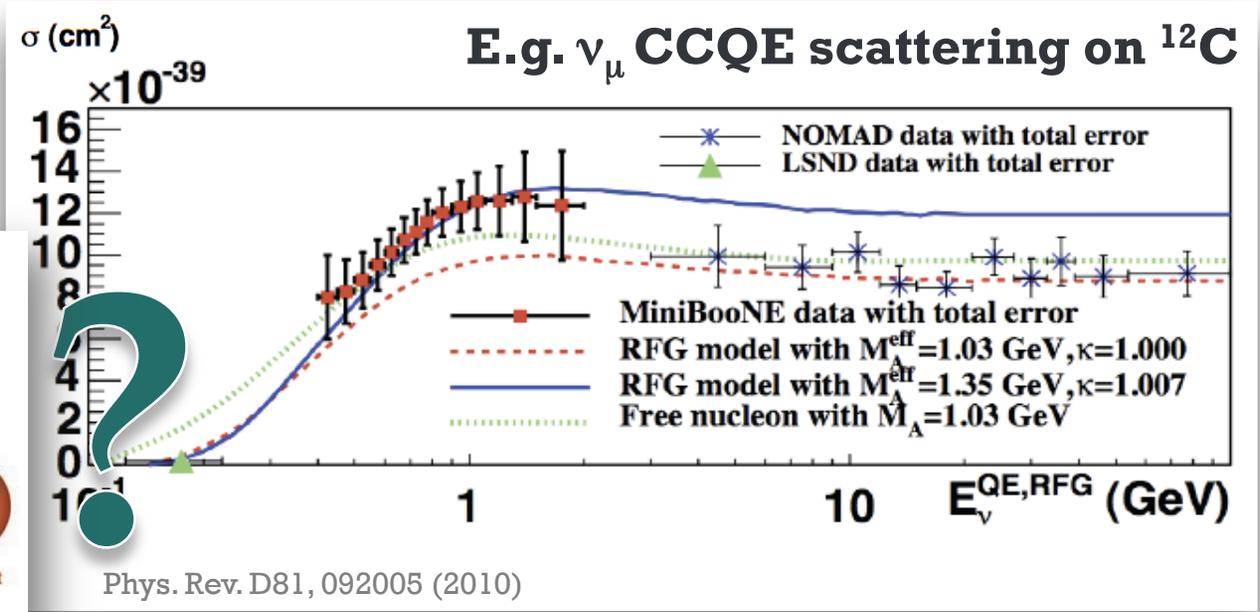
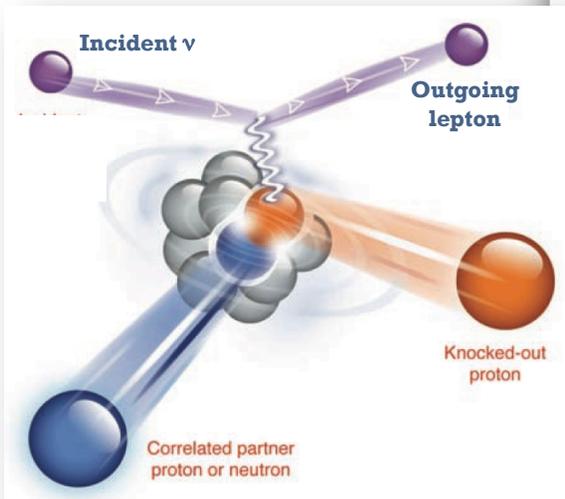


## Additional physics goal

measure neutrino cross sections around 1 GeV

- Past cross section measurements (from K2K, MiniBooNE, SciBooNE, MINOS, NOMAD) have revealed limitations in our understanding of neutrino interactions

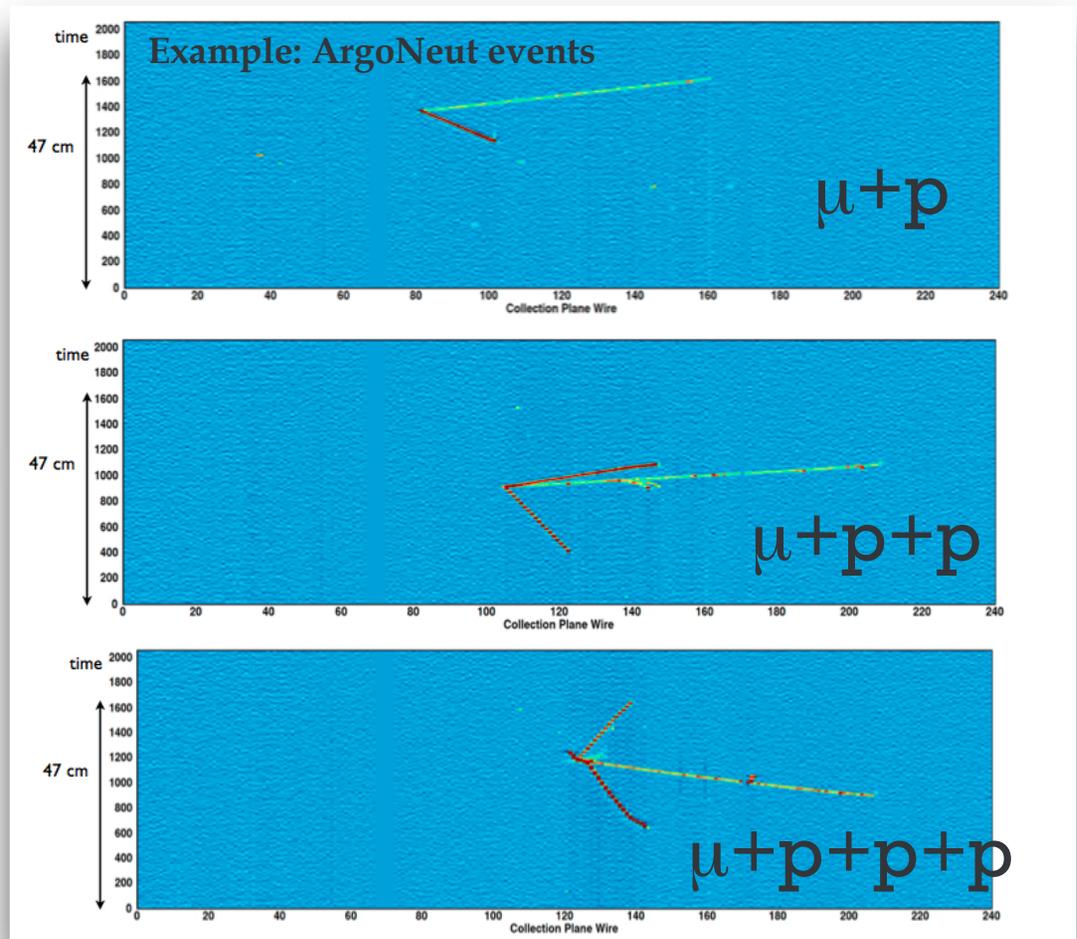
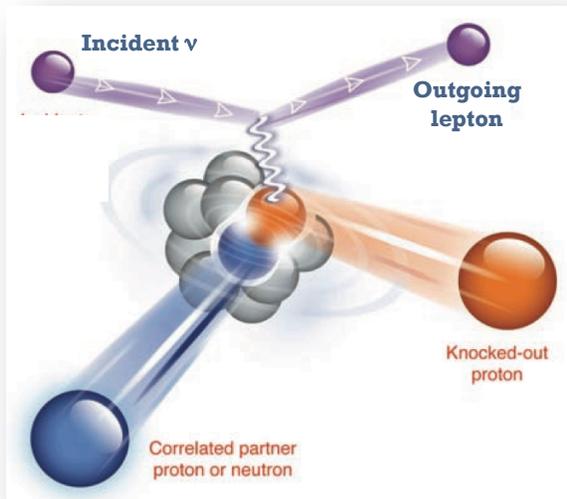
Hadronic effects must play a critical role!



## Additional physics goal

measure neutrino cross sections around 1 GeV

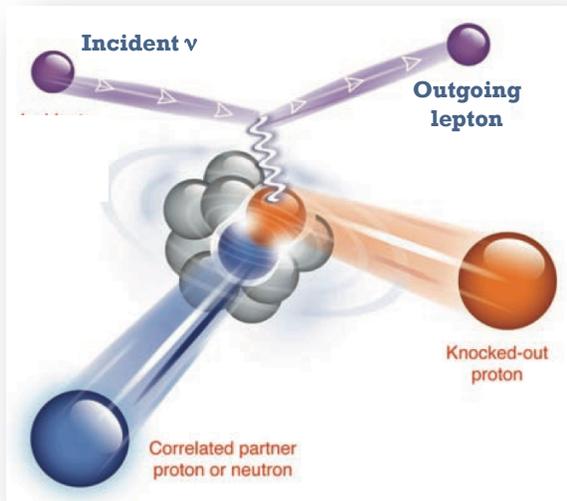
- **Event topology and final state information** (proton multiplicity, momenta) provides information for nuclear effects modeling



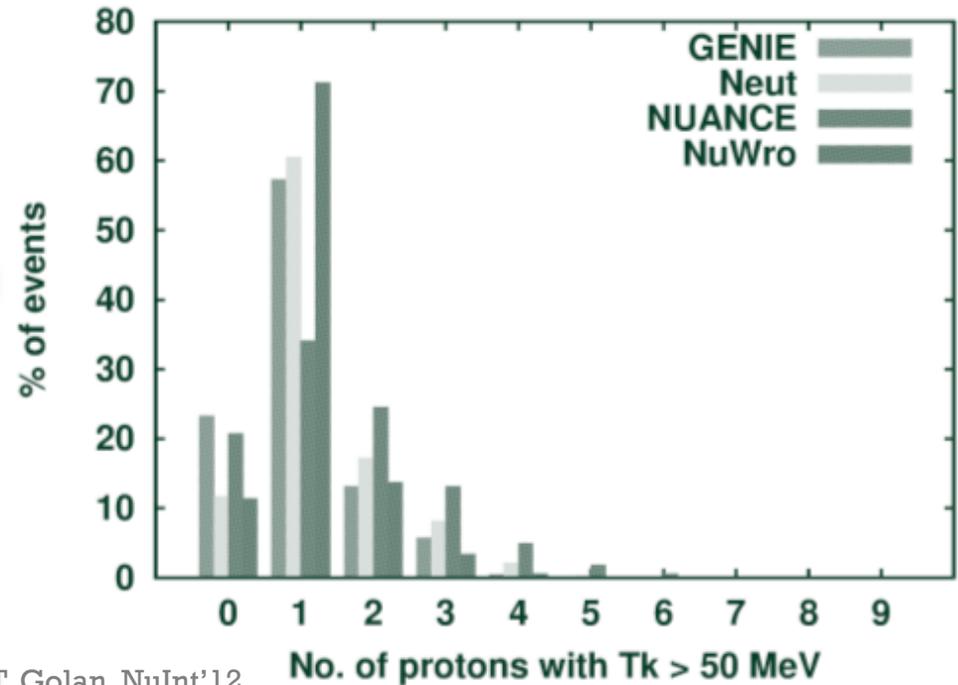
## Additional physics goal

measure neutrino cross sections around 1 GeV

- **Event topology and final state information** (proton multiplicity, momenta) provides information for nuclear effects modeling



Cross-section and nuclear effects modeling: Vastly different predictions from different neutrino event generators

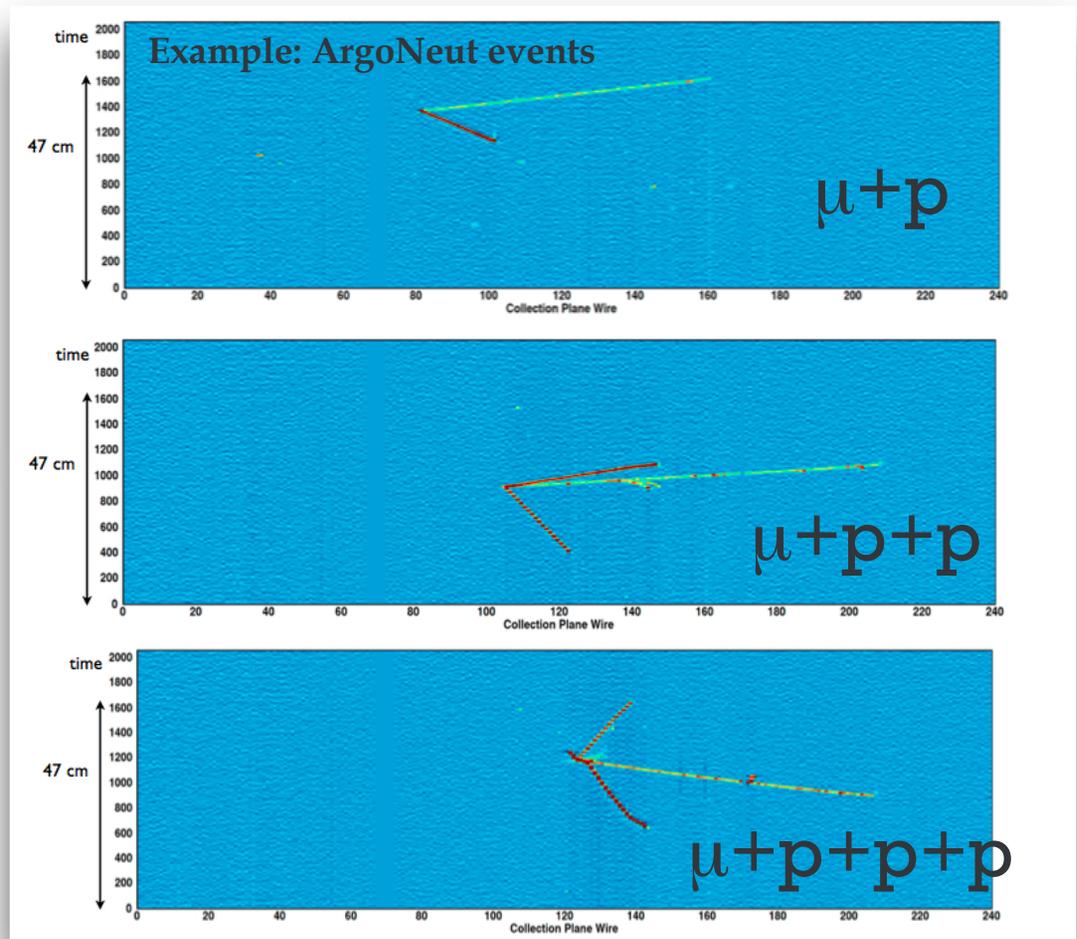
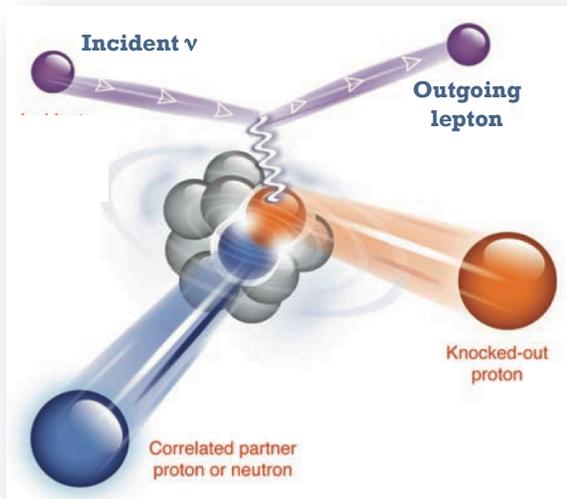


T. Golan, NuInt'12

## Additional physics goal

measure neutrino cross sections around 1 GeV

- **Event topology and final state information** (proton multiplicity, momenta) provides information for nuclear effects modeling



## Additional physics goal

measure neutrino cross sections around 1 GeV

- First LArTPC with high-statistics event samples in 1 GeV range

Nuance-generated events on LAr, MicroBooNE Collaboration

<b>Expected MicroBooNE event rates</b>	
production mode	# events
CC QE ( $\nu_\mu n \rightarrow \mu^- p$ )	60,161
NC elastic ( $\nu_\mu N \rightarrow \nu_\mu N$ )	19,409
CC resonant $\pi^+$ ( $\nu_\mu N \rightarrow \mu^- N \pi^+$ )	25,149
CC resonant $\pi^0$ ( $\nu_\mu n \rightarrow \mu^- p \pi^0$ )	6,994
NC resonant $\pi^0$ ( $\nu_\mu N \rightarrow \nu_\mu N \pi^0$ )	7,388
NC resonant $\pi^\pm$ ( $\nu_\mu N \rightarrow \nu_\mu N' \pi^\pm$ )	4,796
CC DIS ( $\nu_\mu N \rightarrow \mu^- X, W > 2 \text{ GeV}$ )	1,229
NC DIS ( $\nu_\mu N \rightarrow \nu_\mu X, W > 2 \text{ GeV}$ )	456
NC coherent $\pi^0$ ( $\nu_\mu A \rightarrow \nu_\mu A \pi^0$ )	1,694
CC coherent $\pi^+$ ( $\nu_\mu A \rightarrow \mu^- A \pi^+$ )	2,626
NC kaon ( $\nu_\mu N \rightarrow \nu_\mu K X$ )	39
CC kaon ( $\nu_\mu N \rightarrow \mu^- K X$ )	117
other $\nu_\mu$	3,678
total $\nu_\mu$ CC	98,849
total $\nu_\mu$ NC+CC	133,580
$\nu_e$ QE	326
$\nu_e$ CC	657

Preliminary

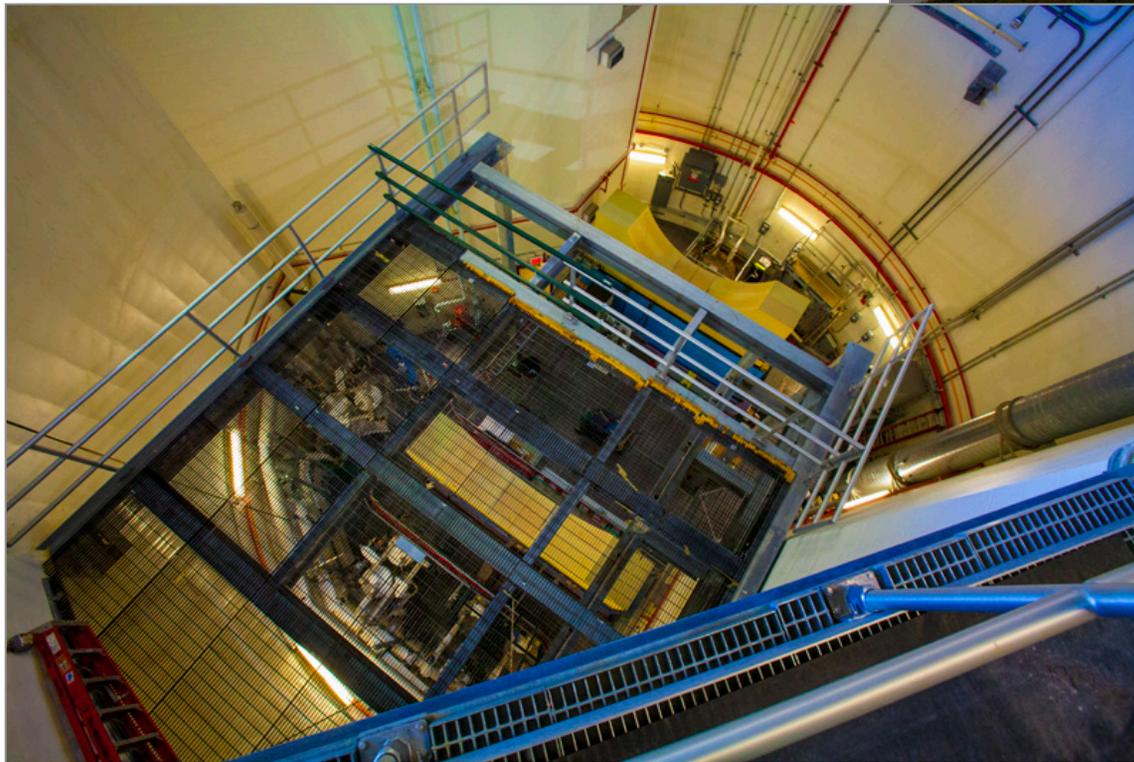
## Additional physics opportunities

- Backgrounds to proton decay & baryon number violating processes for larger (underground) detectors
- Supernova core collapse neutrinos

## Current status

Experiment is well under construction

- LArTF building construction complete, being outfitted for detector for installation



LArTF (Sep. 2013)

# MicroBooNE

60

## Current status

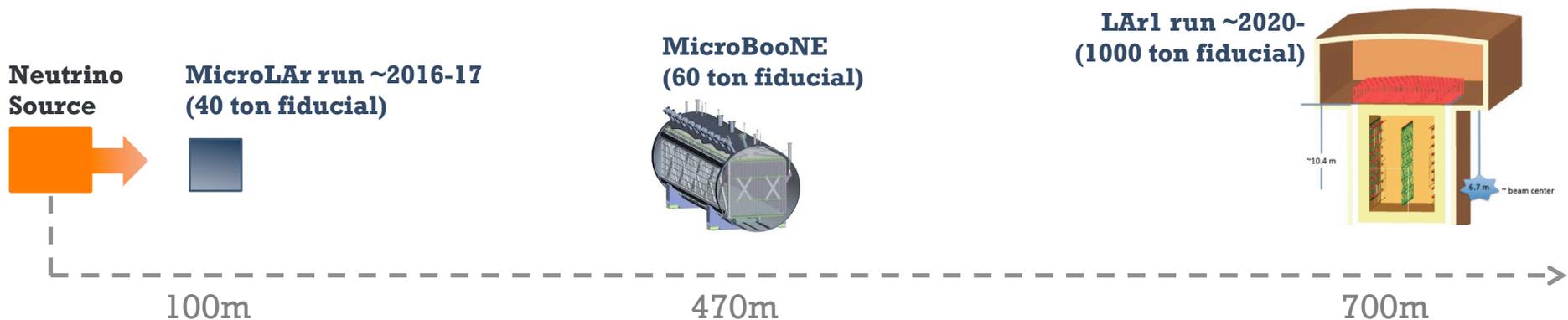
Experiment is well under construction

- LArTF building construction complete, being outfitted for detector for installation
  - PMT's installed in cryostat
  - TPC field cage and wire planes constructed
  - Cold electronics installed on TPC
  - Electronics fully tested and ready for commissioning
- 
- Expected start of data taking: 2014
  - Current MicroBooNE run plan: neutrino mode running,  $6.6e^{20}$  POT (2-3 years to complete)



# Beyond MicroBooNE: LAr1

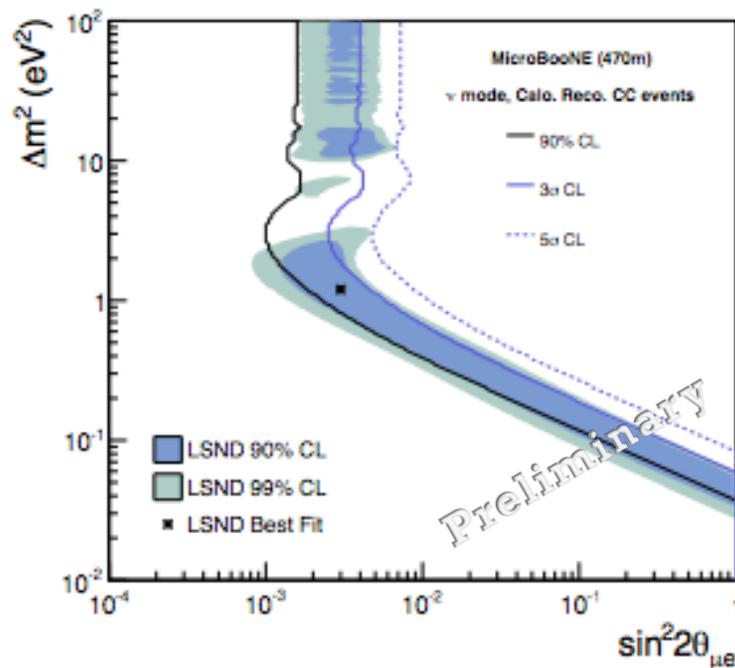
- **New idea:** A second (Phase I) and third (Phase II) LArTPC placed in the Booster Neutrino Beam at Fermilab, in line with MicroBooNE
- Near/far comparison for short-baseline oscillation search
- **Definitive test of MiniBooNE/LSND sterile neutrino interpretation**
- Design philosophy: serve as a development step toward LBNE while functioning as a physics experiment



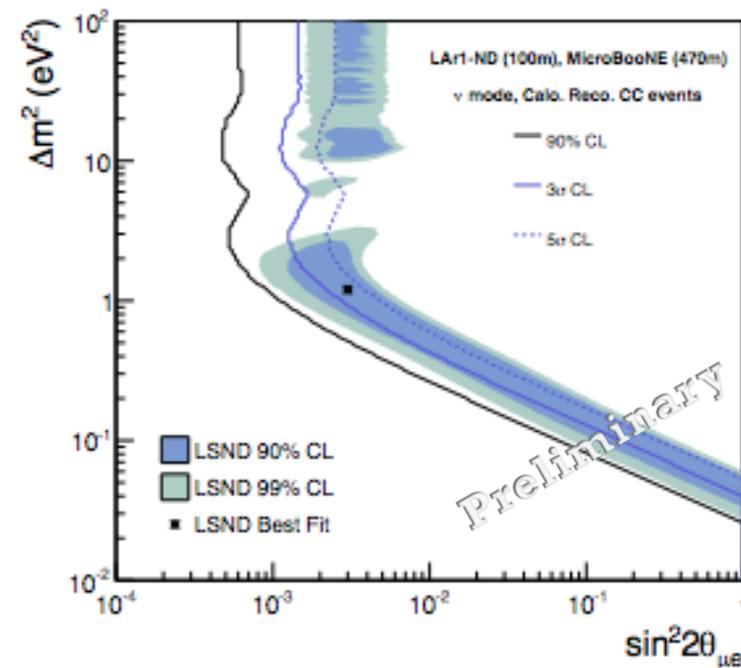
# Beyond MicroBooNE: LAr1 (ND)

The high-statistics event sample in LAr1-ND constrains the expected background event rate in MicroBooNE, reducing the systematic uncertainties.

6.6x10<sup>20</sup> POT exposure for MicroBooNE alone,  
assuming 20% systematic uncertainties  
on  $\nu_e$  backgrounds



Same MicroBooNE exposure +  
2.2x10<sup>20</sup> POT exposure for LAr1-ND



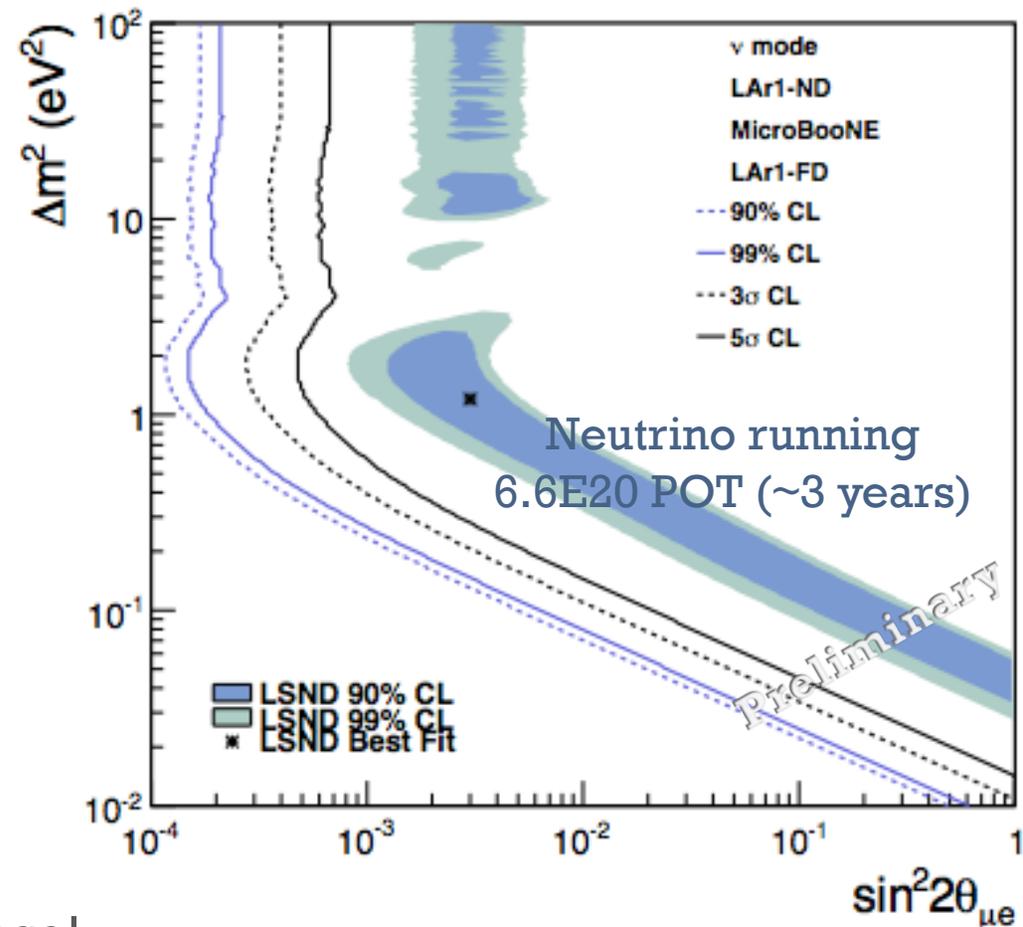
The sensitivity is strengthened through the reduction of systematic errors, covering the LSND best-fit point at  $\sim 4 \sigma$ .

# Beyond MicroBooNE: LAr1

- **Physics reach:** Definitive ( $5\sigma$ ) test of LSND and MiniBooNE in both neutrino and antineutrino modes

While MicroBooNE will definitively address the MiniBooNE excess as electrons or photons in neutrino mode,

the combination of LAr1 (ND) and MicroBooNE will allow for its interpretation as new physics.



- Also  $(\bar{\nu}_e)$  and  $(\bar{\nu}_\mu)$  disappearance!

# Beyond MicroBooNE: LAr1

## Current status

- Letter of Intent submitted to Fermilab Directorate in 2012
- Snowmass White Paper in Sep. 2013, focused on LAr1-ND and potential to expand into LAr1-ND (arXiv:1309.7987)
- Strong ongoing effort to develop full LAr1-ND proposal by Dec. 2013

A Letter of Intent for a Neutrino Oscillation Experiment on the  
Booster Neutrino Beamline: LAr1

June 13, 2012

H. Chen, C. Thorn, D. Lissauer, V. Radeka, B. Yu, G. Mahler, S. Rescia, S. Duffin, Y. Li  
*Brookhaven National Laboratory, Upton, NY*

L. Bartoszek  
*Bartoszek Engineering*

E. Blucher, D. Schmitz  
*University of Chicago, Chicago, IL*

D. Kaleko, G. Karagiorgi, B. Seligman, M. Shaevitz, B. Willis  
*Columbia University, Nevis Labs, Irvington, NY 10533*

B. Baller, H. Greenlee, J. Raaf, R. Rameika, G. Zeller  
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M. Messier, S. Mufson, J. Musser, J. Urheim  
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W. Huelsnitz, W. C. Louis, G. B. Mills, Z. Pavlovic, R. G. Van De Water  
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L. Bugel, J. Conrad, T. Katori, C. Ignarra, B. Jones, M. Toups  
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*Syracuse University, Syracuse, NY*

M. Marshak  
*University of Minnesota, Minneapolis, MN, 55455*

F. Cavanna, E. Church, B. Fleming, R. Guenette, O. Palamara, K. Partyka, A. Szelc  
*Yale University, New Haven, CT 06520*

# Summary

We are entering an era of **high-precision neutrino physics** measurements, enabled by the **LArTPC** technology:

- Increased detection efficiency (>x2 conventional detectors)
- Increased background rejection from detailed event topology and  $dE/dx$

**MicroBooNE** is first in the line of ton-scale LArTPC detectors in the US.

- Important R&D stepping stone for LBNE (mass hierarchy, CP violation,...)
- Extremely **rich physics** on its own:
  - MiniBooNE low energy excess
  - Neutrino cross sections on Ar
  - Sterile neutrino oscillations with the addition of second/third detector
  - ...
- MicroBooNE data taking begins in less than a year!

# End remarks

Theoretical motivation for light ( $\sim 1$  eV) sterile neutrinos is perhaps not so strong, though sterile neutrinos with sizable mixing emerge in several models of neutrino mass (heavy sterile neutrinos...).

Their discovery would point towards new physics.

“...their role is relevant enough to justify an open mind attitude and a close look for any, yet tiny, evidence for new effects beyond the *too much* successful Standard Model.”

[Theorist Anonymous]

Experimental hints may be right in front of us, albeit not completely understood. Need new, definitive experiments. Model-independent searches should be given highest priority.

Thank you!